Einstein Metrics,

Four-Manifolds, &

Gravitational Instantons

Claude LeBrun Stony Brook University

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Int. Math. Res. Not. IMRN

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Will also briefly mention

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with

Tristan Ozuch MIT

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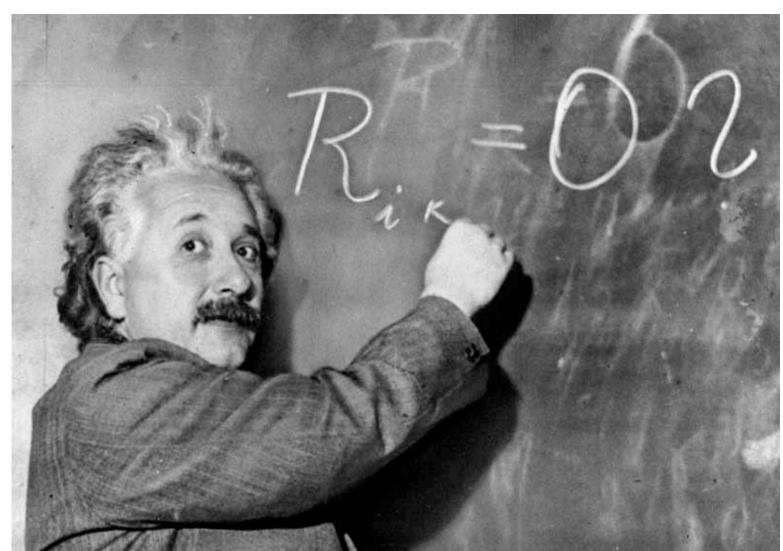
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"...the greatest blunder of my life!"

— A. Einstein, to G. Gamow

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As punishment ...

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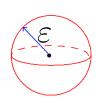
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$$s = r_j^j = \mathcal{R}^{ij}{}_{ij}.$$

$$\frac{\operatorname{vol}_g(B_{\varepsilon}(p))}{c_n \varepsilon^n} = 1 - s \frac{\varepsilon^2}{6(n+2)} + O(\varepsilon^4)$$



Einstein's equations are "locally trivial:"

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⇒ Global rigidity results in these low dimensions.

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(Catanese-LeBrun, Rasdeaconu-Şuvaina)

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In high dimensions, the Einstein condition allows for a surprising degree of flexibility!

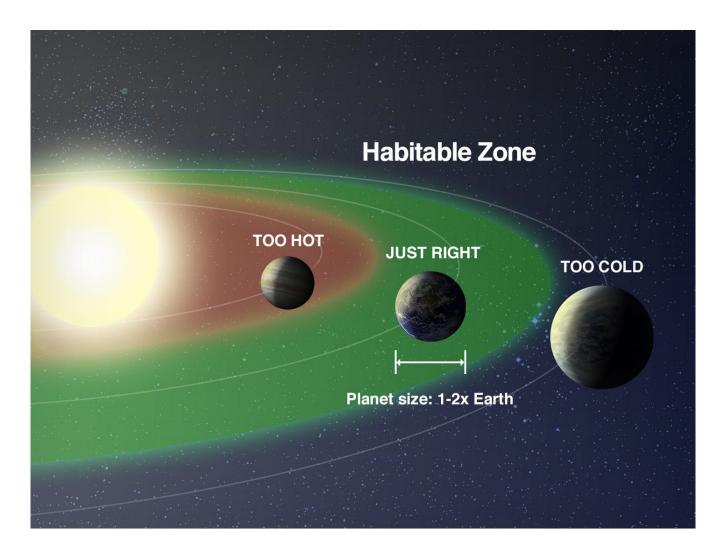
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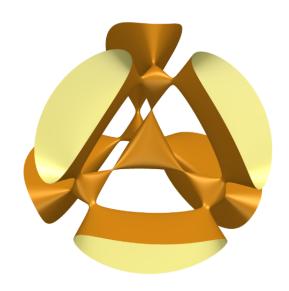
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Theorem (LeBrun). There is only one Einstein metric on compact complex-hyperbolic 4-manifold $\mathbb{C}\mathcal{H}_2/\Gamma$, up to scale and diffeos.

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$$\star^2 = 1.$$

 Λ^+ self-dual 2-forms.

 Λ^- anti-self-dual 2-forms.

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Which 4-manifolds admit Einstein metrics?

A laboratory for exploring Einstein metrics.

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$$d\omega = 0, \qquad \exists \omega : TM \stackrel{\cong}{\to} T^*M.$$

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$$\omega = dx \wedge dy + dz \wedge dt$$

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Some Suggestive Questions. If (M^4, ω) is a symplectic 4-manifold, when does M^4 admit an Einstein metric g (a priori unrelated to ω)?

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Some Suggestive Questions. If (M^4, ω) is a symplectic 4-manifold, when does M^4 admit an Einstein metric g (a priori unrelated to ω)? What if we also require $\lambda > 0$?

Theorem (CLW '08). Suppose that M is a smooth compact oriented 4-manifold which carries some symplectic form ω .

$$\iff M \stackrel{\mathit{diff}}{pprox}$$

$$\iff M \stackrel{diff}{pprox} \left\{ \begin{array}{c} \mathbb{CP}_2 \# k \overline{\mathbb{CP}}_2, \end{array} \right.$$

$$\iff M \stackrel{\text{diff}}{\approx} \left\{ \begin{array}{c} \mathbb{CP}_2 \# k \overline{\mathbb{CP}}_2, & 0 \le k \le 8, \\ \end{array} \right.$$

$$\iff M \stackrel{\text{diff}}{\approx} \begin{cases} \mathbb{CP}_2 \# k \overline{\mathbb{CP}}_2, & 0 \le k \le 8, \\ or \\ S^2 \times S^2 \end{cases}$$

 $\overline{\mathbb{CP}}_2$ = reverse oriented \mathbb{CP}_2 .

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Diffeotypes: exactly the Del Pezzo surfaces.

 (M^4, J) for which c_1 is a Kähler class $[\omega]$.

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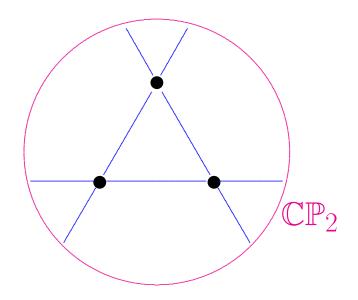
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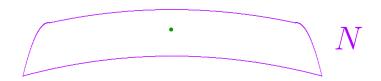
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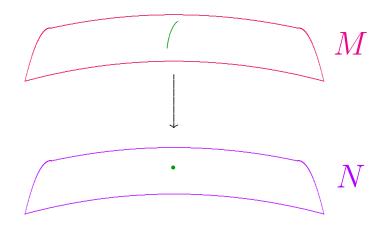
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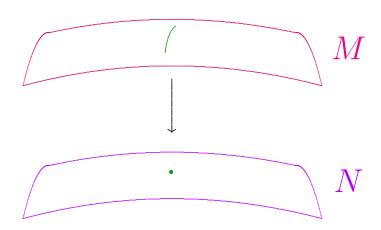


If N is a complex surface, may replace $p \in N$ with \mathbb{CP}_1



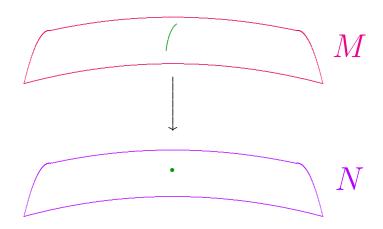
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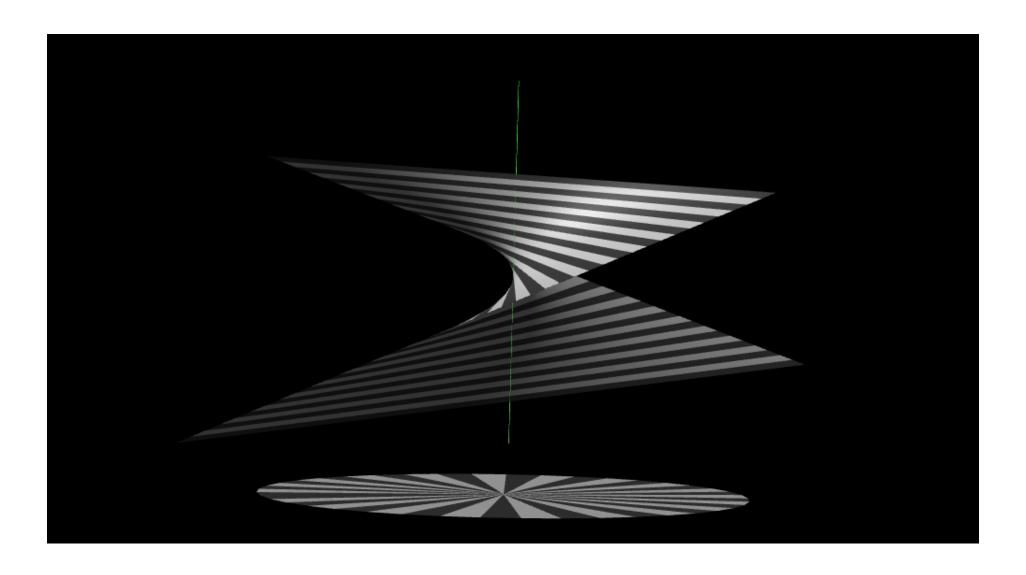
$$M \approx N \# \overline{\mathbb{CP}}_2$$



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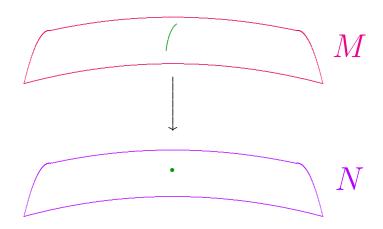
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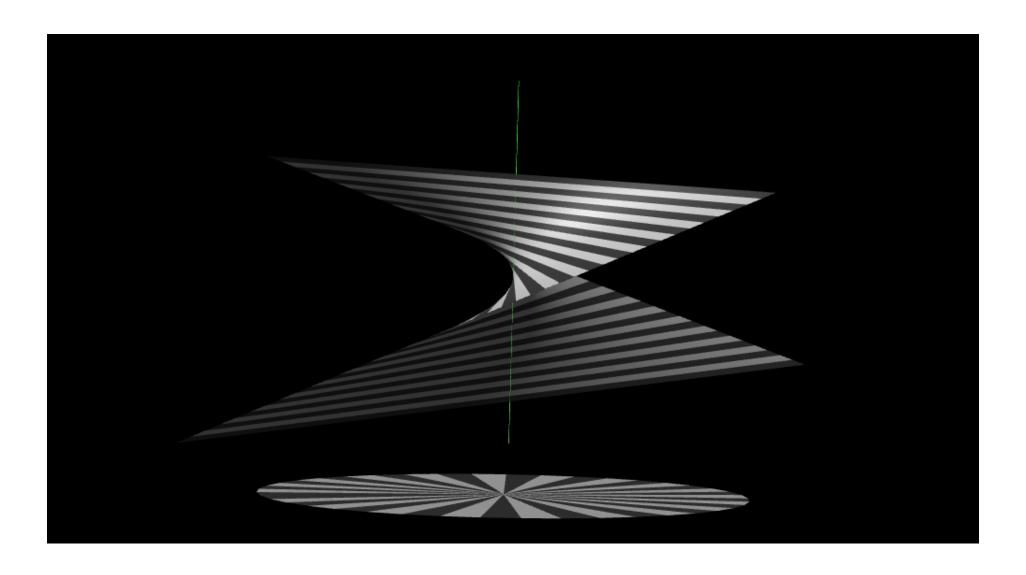




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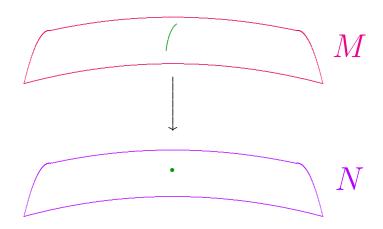
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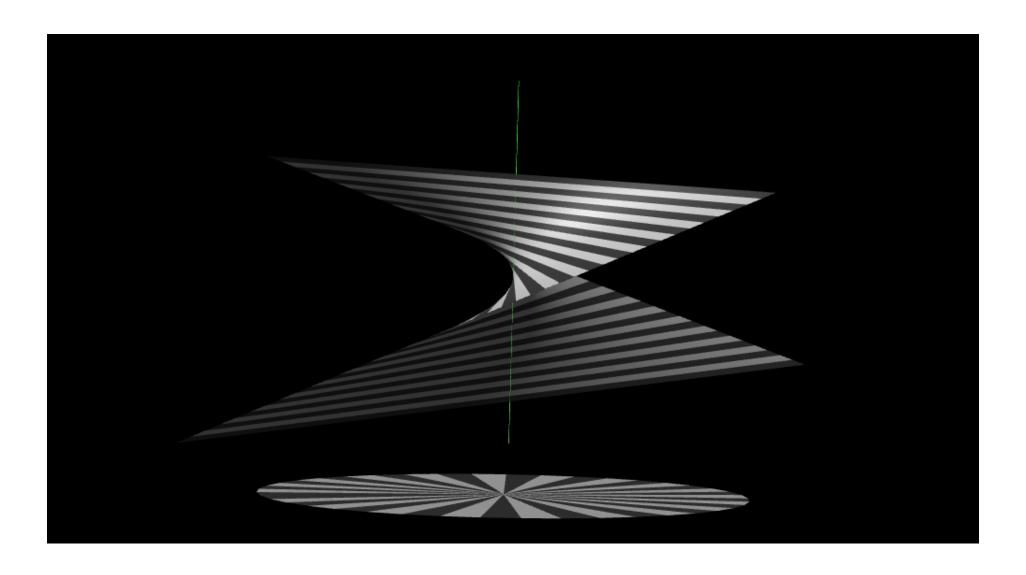




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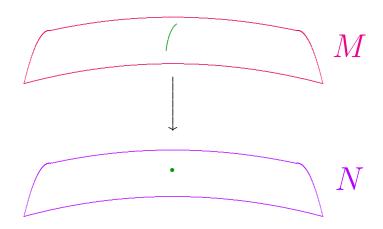
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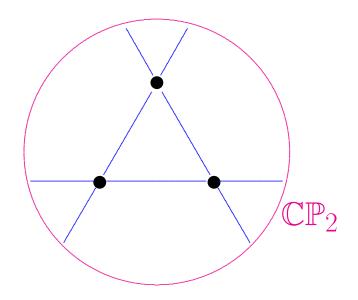
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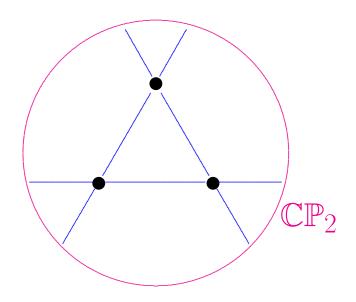
 (M^4, J) for which c_1 is a Kähler class $[\omega]$. Shorthand: " $c_1 > 0$."

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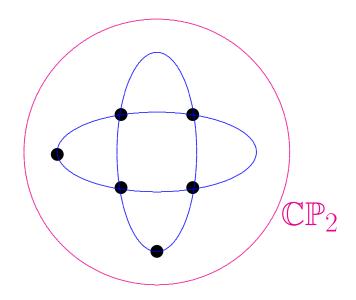
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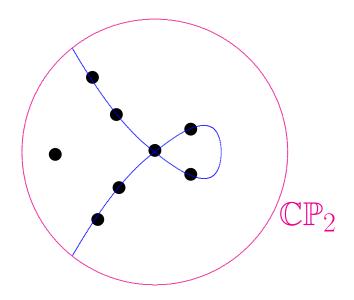
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No 3 on a line, no 6 on conic, no 8 on nodal cubic.

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$$g = g(J \cdot, J \cdot)$$

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Theorem. Each del Pezzo (M^4, J) admits a J-compatible Hermitian, Einstein metric, and this metric is unique up to complex automorphisms and constant rescalings.

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Top eigenspace $L \subset \Lambda^+$ of W_+ is a line bundle.

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at every point, with respect to h. Now integrate!

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Corollary. Every simply-connected compact oriented Einstein (M^4, h) with $det(W_+) > 0$ is diffeomorphic to a del Pezzo surface.

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Corollary. Every simply-connected compact oriented Einstein (M^4, h) with $det(W_+) > 0$ is diffeomorphic to a del Pezzo surface. Conversely, every del Pezzo M^4 carries Einstein h with $det(W_+) > 0$, and these sweep out exactly one connected component of moduli space $\mathscr{E}(M)$.

Objective: if orbifold limit is conformally Kähler, show that the same is also true of smooth 4-manifolds far out in the sequence.

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This illustrates how gravitational instantons play a crucial role, even when studying compact case.

Gravitational Instantons?

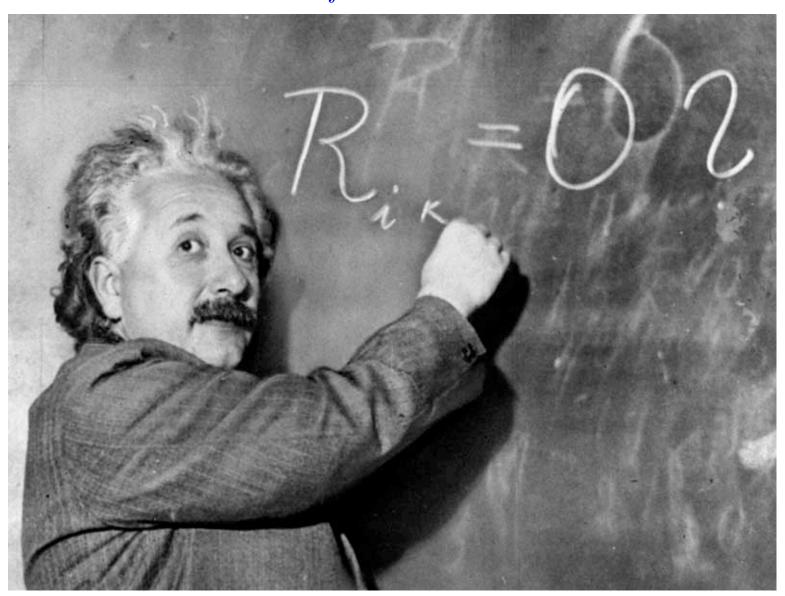
Definition. A gravitational instanton is a

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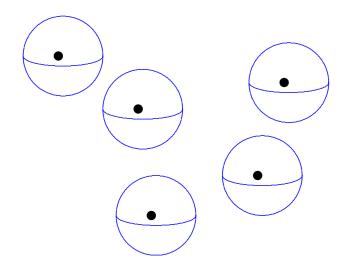
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Data: ℓ points in \mathbb{R}^3 and $\kappa^2 \Longrightarrow V$ with $\Delta V = 0$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\varrho_j}$$

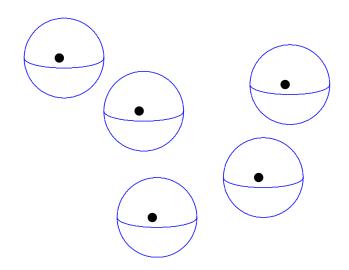
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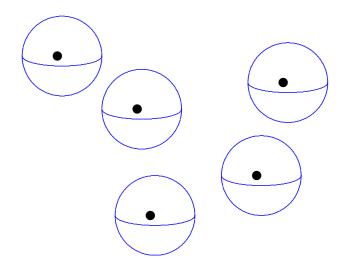
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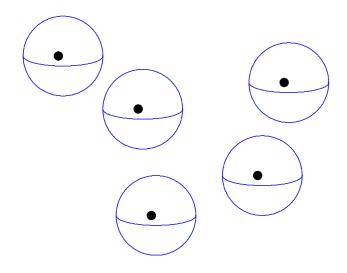
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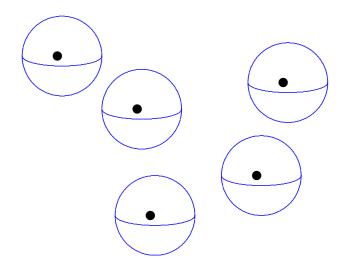
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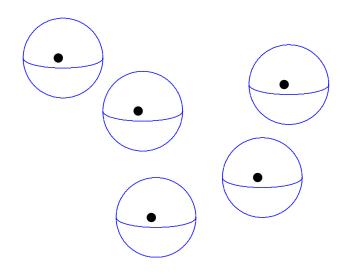
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Gibbons and Hawking were unaware of all this!

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The $\ell = 2$ case is Eguchi-Hanson $\approx T^*S^2$.

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This last property distinguishes the ALF spaces from other classes of gravitational instantons:

ALG, ALH, ALG*, ALH*, ...

Example.

$$g = V(dx^{2} + dy^{2} + dz^{2}) + V^{-1}\theta^{2}$$
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for left-invariant coframe $\{\sigma_j\}$ on $S^3 = \mathbf{SU}(2)$.

Taub-NUT becomes Hermitian metric on \mathbb{C}^2 .

Non-Kähler, but conformally Kähler!

Hawking also explored non-hyper-Kähler examples...

Example.

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

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Conformal to

$$\mathbf{h} = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho} \right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho} \right) dt^2 \right] + g_{S^2}$$

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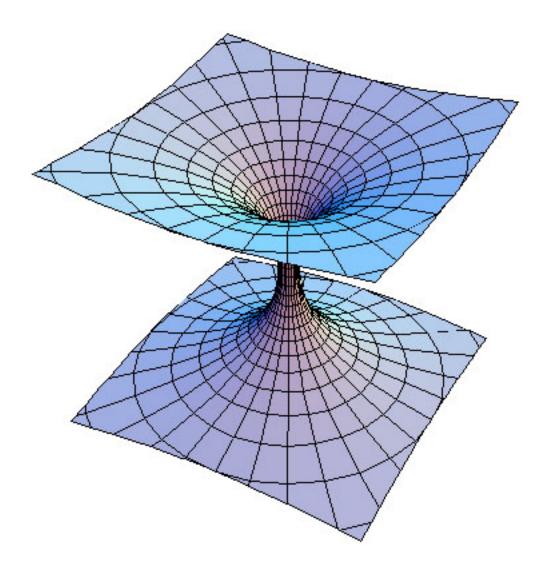
$$g = dr^2 + r^2 d\theta^2 + 4m^2 g_{S^2} + O(r^2)$$

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Hawking: set $t = 4m\theta$ and $\varrho = 2m + \frac{r^2}{8m}$. This makes g into a Ricci-flat metric on $\mathbb{R}^2 \times S^2$. Makes h into extremal Kähler metric on $\mathbb{C} \times \mathbb{CP}_1$.



$$\mathbb{R} \times S^2 \subset \mathbb{R}^2 \times S^2$$

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Hitchin, Kronheimer, Cherkis-Hitchin, Minerbe, Hein, Chen-Chen, Hein-Sun-Viaclovsky-Zhang...

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But my collaborators Biquard and Gauduchon have fortunately done us all the favor of reminding us that the hyper-Kähler gravitons are only one small part of the story!

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Theorem (Biquard-Gauduchon '23). Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J. Also assume that (M, g, J) is not Kähler.

• the (reverse-oriented) Taub-NUT metric;

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Diffeomorphic to $\mathbb{CP}_2 - \{pt\}$

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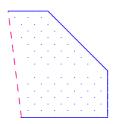
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Y. Chen & E. Teo, 2011

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allowing one to generalize methods first explored in the compact case. Theorem A (BGL '24).

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$$|\mho|_{g_0} = O(\varrho^{-1}), \quad |\nabla \mho|_{g_0} = O(\varrho^{-2}), \quad \dots$$

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Theorem A (BGL '24). Let (M, g_0) be any of the ALF toric Hermitian gravitational instantons appearing in the Biquard-Gauduchon classification. Then any other Ricci-flat Riemannian metric g on M which is sufficiently C_1^3 -close to g is conformal to some strictly extremal Kähler metric h, and so is, in particular, Hermitian. Moreover, every such g carries at least one Killing field.

Set $h = \alpha^{2/3}g$, where α top eigenvalue of W_{+g} , and choose top eigenform $\omega \in \Lambda^+$ with $|\omega|_h \equiv 1$. Then

$$0 \ge |\nabla \omega|^2 + 3\langle \omega, (d+d^*)^2 \omega \rangle$$

at every point, with respect to h.

Theorem B (BGL '24). Let (M, g_0) be any toric Hermitian ALF gravitational instanton.

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This optimal result combines **Theorem A** with a result of Mingyang Li, arXiv:2310.13197.

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