Einstein Metrics,

Four-Manifolds, &

Gravitational Instantons

Claude LeBrun Stony Brook University

Complex Hermitian Geometry, Colloque en l'honneur de Paul Gauduchon, Angers, France, 22 mai, 2025.



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with

Tristan Ozuch MIT

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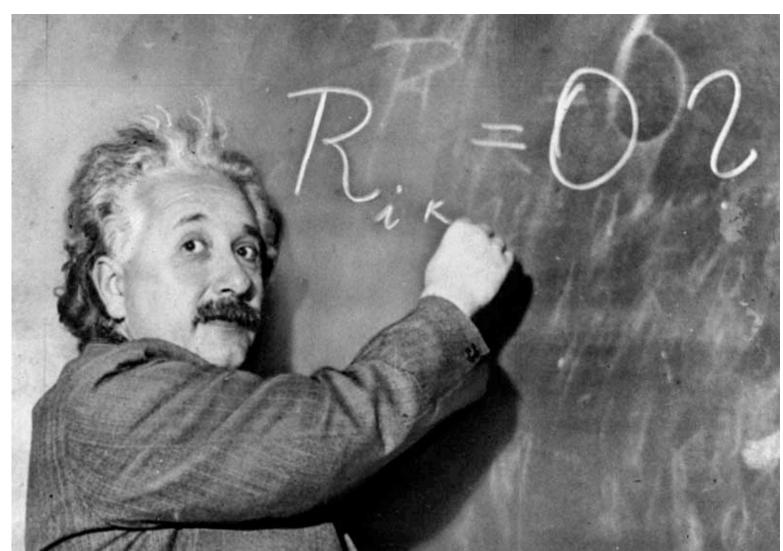
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"... the greatest blunder of my life!"

— A. Einstein, to G. Gamow

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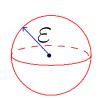
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$$\frac{\operatorname{vol}_g(B_{\varepsilon}(p))}{c_n \varepsilon^n} = 1 - s \frac{\varepsilon^2}{6(n+2)} + O(\varepsilon^4)$$



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⇒ Global rigidity results in these low dimensions.

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(Catanese-LeBrun, Răsdeaconu-Şuvaina)

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In high dimensions, the Einstein condition allows for a surprising degree of flexibility!

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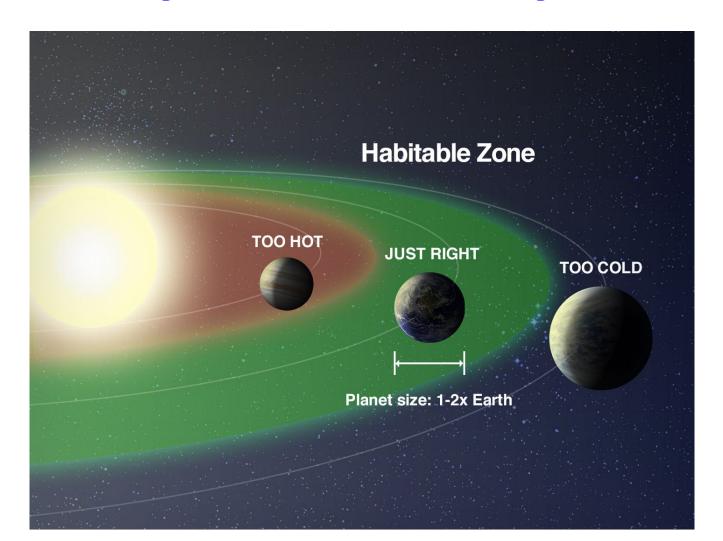
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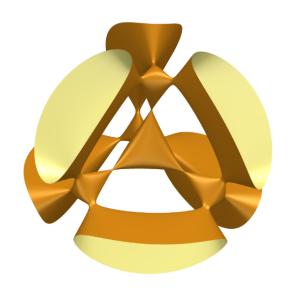
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 $\Lambda^+$  self-dual 2-forms.

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Which 4-manifolds admit Einstein metrics?

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$$d\omega = 0, \qquad \exists \omega : TM \stackrel{\cong}{\to} T^*M.$$

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$$\omega = dx \wedge dy + dz \wedge dt$$

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Some Suggestive Questions. If  $(M^4, \omega)$  is a symplectic 4-manifold, when does  $M^4$  admit an Einstein metric g (a priori unrelated to  $\omega$ )?

## Symplectic 4-manifolds:

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Some Suggestive Questions. If  $(M^4, \omega)$  is a symplectic 4-manifold, when does  $M^4$  admit an Einstein metric g (a priori unrelated to  $\omega$ )? What if we also require  $\lambda > 0$ ?

Theorem (CLW '08). Suppose that M is a smooth compact oriented 4-manifold which carries some symplectic form  $\omega$ .

$$\iff M \stackrel{\mathit{diff}}{pprox}$$

$$\iff M \stackrel{diff}{pprox} \left\{ \begin{array}{c} \mathbb{CP}_2 \# k \overline{\mathbb{CP}}_2, \end{array} \right.$$

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 $\overline{\mathbb{CP}}_2$  = reverse oriented  $\mathbb{CP}_2$ .

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Diffeotypes: exactly the Del Pezzo surfaces.

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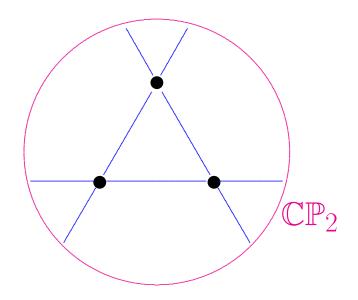
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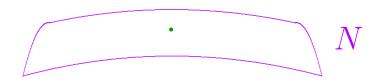
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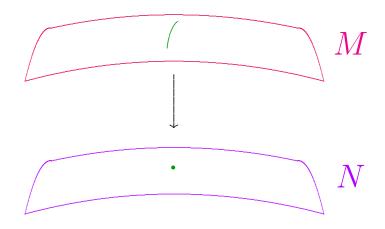
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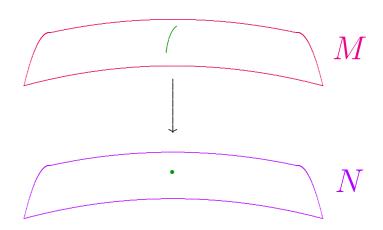


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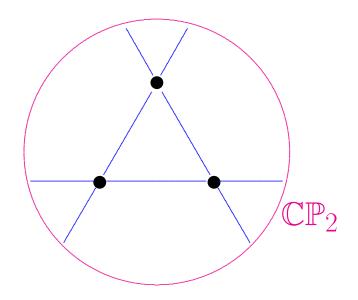
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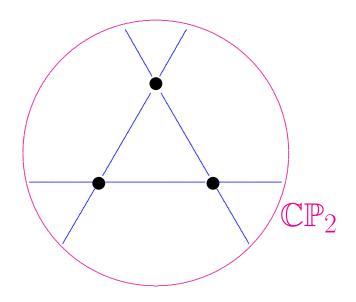
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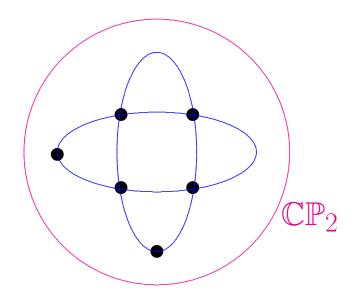
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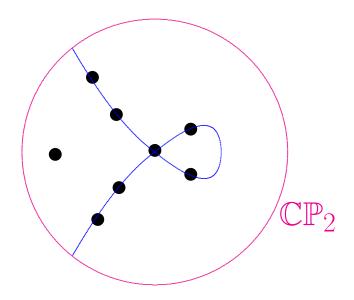
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No 3 on a line, no 6 on conic, no 8 on nodal cubic.

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Theorem.

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**Theorem.** Each del Pezzo  $(M^4, J)$  admits a J-compatible

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$$g = g(J \cdot, J \cdot)$$

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**Theorem.** Each del Pezzo  $(M^4, J)$  admits a J-compatible Hermitian, Einstein metric, and this metric is unique up to complex automorphisms and constant rescalings.

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L '97: Such Einstein metrics are necessarily conformal to extremal Kähler metrics, in the sense of Calabi.

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 $(M^4, J)$  for which  $c_1$  is a Kähler class  $[\omega]$ . Shorthand: " $c_1 > 0$ ."

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L '15 First successful approach, in terms of properties of unique self-dual harmonic 2-form.

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Top eigenspace  $L \subset \Lambda^+$  of  $W_+$  is a line bundle.

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at every point, with respect to h.

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at every point, with respect to h. Now integrate!

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Corollary. Every simply-connected compact oriented Einstein  $(M^4, h)$  with  $det(W_+) > 0$  is diffeomorphic to a del Pezzo surface.

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Corollary. Every simply-connected compact oriented Einstein  $(M^4, h)$  with  $det(W_+) > 0$  is diffeomorphic to a del Pezzo surface. Conversely, every del Pezzo  $M^4$  carries Einstein h with  $det(W_+) > 0$ , and these sweep out exactly one connected component of moduli space  $\mathscr{E}(M)$ .

Objective: if orbifold limit is conformally Kähler, show that the same is also true of smooth 4-manifolds far out in the sequence.

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This illustrates how gravitational instantons play a crucial role, even when studying compact case.

## **Gravitational Instantons?**

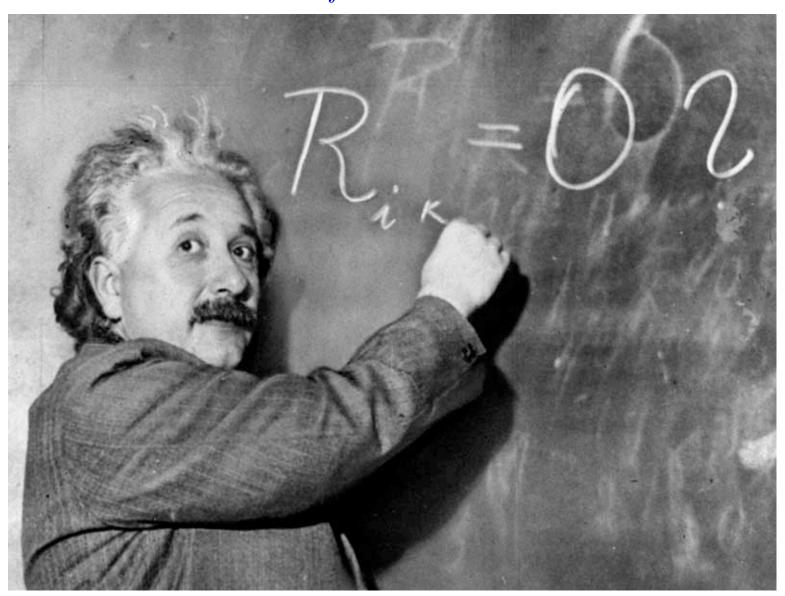
Definition. A gravitational instanton is a

Definition. A gravitational instanton is a complete,

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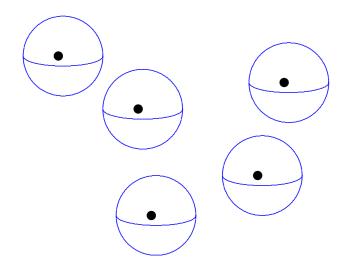
•

Data:  $\ell$  points in  $\mathbb{R}^3$  and  $\kappa^2 \Longrightarrow V$  with  $\Delta V = 0$ 

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\varrho_j}$$

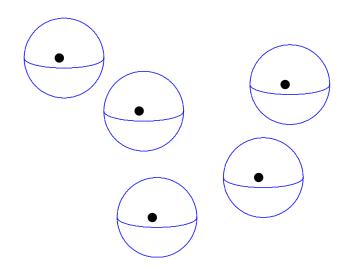
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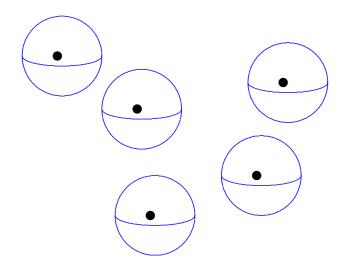
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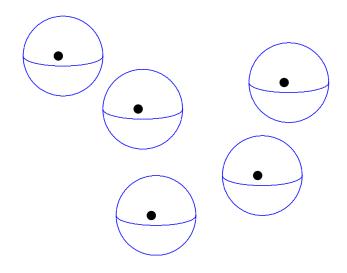
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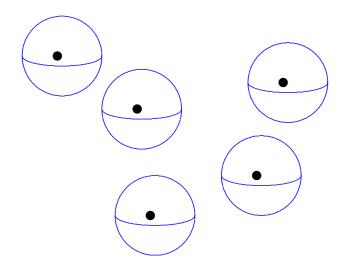
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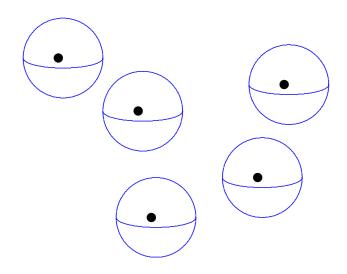
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 $M \to \mathbb{R}^3$  hyper-Kähler moment map of  $S^1$  action.

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Gibbons and Hawking were unaware of all this!

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The  $\ell = 2$  case is Eguchi-Hanson  $\approx T^*S^2$ .

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This last property distinguishes the ALF spaces from other classes of gravitational instantons:

ALG, ALH, ALG\*, ALH\*, ...

# Example.

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$$g = \frac{r+1}{4r}dr^2 + r(1+r)[\sigma_1^2 + \sigma_2^2] + \frac{r}{r+1}\sigma_3^2$$

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for left-invariant coframe  $\{\sigma_j\}$  on  $S^3 = \mathbf{SU}(2)$ .

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$$g = \frac{r+1}{4r}dr^2 + r(1+r)[\sigma_1^2 + \sigma_2^2] + \frac{r}{r+1}\sigma_3^2$$

for left-invariant coframe  $\{\sigma_j\}$  on  $S^3 = \mathbf{SU}(2)$ .

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This J determines opposite orientation from the hyper-Kähler complex structures.

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Non-Kähler, but conformally Kähler!

Hawking also explored non-hyper-Kähler examples...

# Example.

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

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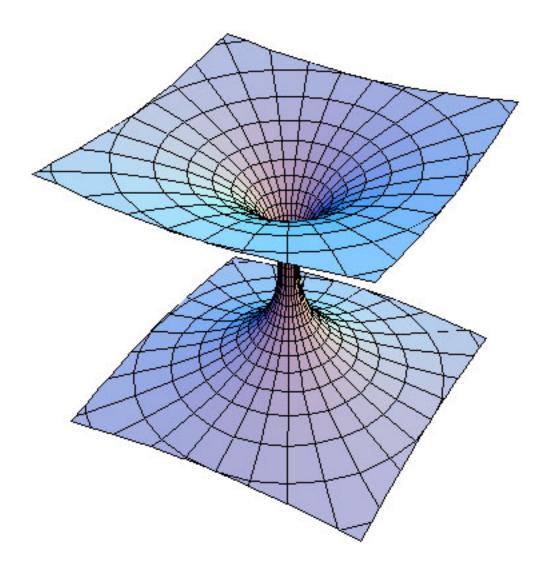
$$g = dr^2 + r^2 d\theta^2 + 4m^2 g_{S^2} + O(r^2)$$

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$$\mathbb{R} \times S^2 \subset \mathbb{R}^2 \times S^2$$

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Hitchin, Kronheimer, Cherkis-Hitchin, Minerbe, Hein, Chen-Chen, Hein-Sun-Viaclovsky-Zhang...

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This might lend some credence to the aphorism...

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"Die Mathematiker sind eine Art Franzosen: redet man zu ihnen, so übersetzen sie es in ihre Sprache, und dann ist es alsobald ganz etwas anderes."

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"Mathematicians are like Frenchmen: you tell them something, they translate it into their own language, and before you know it, it's something else entirely."

— J.W. von Goethe

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But now my French collaborators Biquard and Gauduchon have fortunately done us all the favor of reminding us that the hyper-Kähler gravitons are only one small part of the story!

**Theorem** (Biquard-Gauduchon '23). Let  $(M^4, g)$  be a smooth, complete, non-flat,

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 $\mathbb{T}^2$  acts effectively and isometrically

$$g = d\varrho^2 + \varrho^2 \gamma + \eta^2 + \mho$$

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$$\nabla = O(\varrho^{-1}), \quad \nabla \nabla = O(\varrho^{-2}), \quad \dots \quad \nabla^3 \nabla = O(\varrho^{-4})$$

$$\implies \operatorname{Vol}(B_{\rho}) \sim \operatorname{const} \cdot \rho^3$$

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$$g(J\cdot, J\cdot) = g$$

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• the (reverse-oriented) Taub-NUT metric;

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Diffeomorphic to  $\mathbb{CP}_2 - \{pt\}$ 

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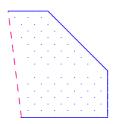
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Y. Chen & E. Teo, 2011

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**Theorem** (Biquard-Gauduchon '23). Let  $(M^4, g)$  be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J. Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:

- the (reverse-oriented) Taub-NUT metric;
- the Taub-bolt metric;
- a metric of the Kerr family; or
- a metric in the Chen-Teo family.

This implies that they always satisfy Peng Wu's criterion

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Set  $h = \alpha^{2/3}g$ , where  $\alpha$  top eigenvalue of  $W_{+g}$ , and choose top eigenform  $\omega \in \Lambda^+$  with  $|\omega|_h \equiv 1$ . Then

$$0 \ge |\nabla \omega|^2 + 3\langle \omega, (d+d^*)^2 \omega \rangle$$

at every point, with respect to h.

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at every point, with respect to h. Now use Stokes!

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The metric h also satisfies

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With appropriate fall-off at infinity, this can be used to control the boundary term.

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$$|\mho|_{g_0} = O(\varrho^{-1}), \quad |\nabla \mho|_{g_0} = O(\varrho^{-2}), \quad \dots$$

**Theorem A** (BGL '24). Let  $(M, g_0)$  be any of the ALF toric Hermitian gravitational instantons appearing in the Biquard-Gauduchon classification. Then any other Ricci-flat Riemannian metric g on M which is sufficiently  $C_1^3$ -close to g is conformal to some strictly extremal Kähler metric h,

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This optimal result combines **Theorem A** with a result of Mingyang Li, arXiv:2310.13197.

### Merci de m'avoir invité!

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### Je suis ravi d'être ici!



# Félicitations, Paul!



## Et Bon Anniversaire!

