

Einstein Manifolds,
Self-Dual Weyl Weyl Curvature, &
Conformally Kähler Geometry

Claude LeBrun
Stony Brook University

Interactions in Complex Geometry
Vanderbilt University, December 12, 2021

Most recent references:

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Mathematical Research Letters
28 (2021) 127–144

And

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Einstein Manifolds, Conformal Curvature, and
Anti-Holomorphic Involutions

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Annales Mathématiques du Québec
45(2) (2021) 391–405

Definition. *A Riemannian metric h*

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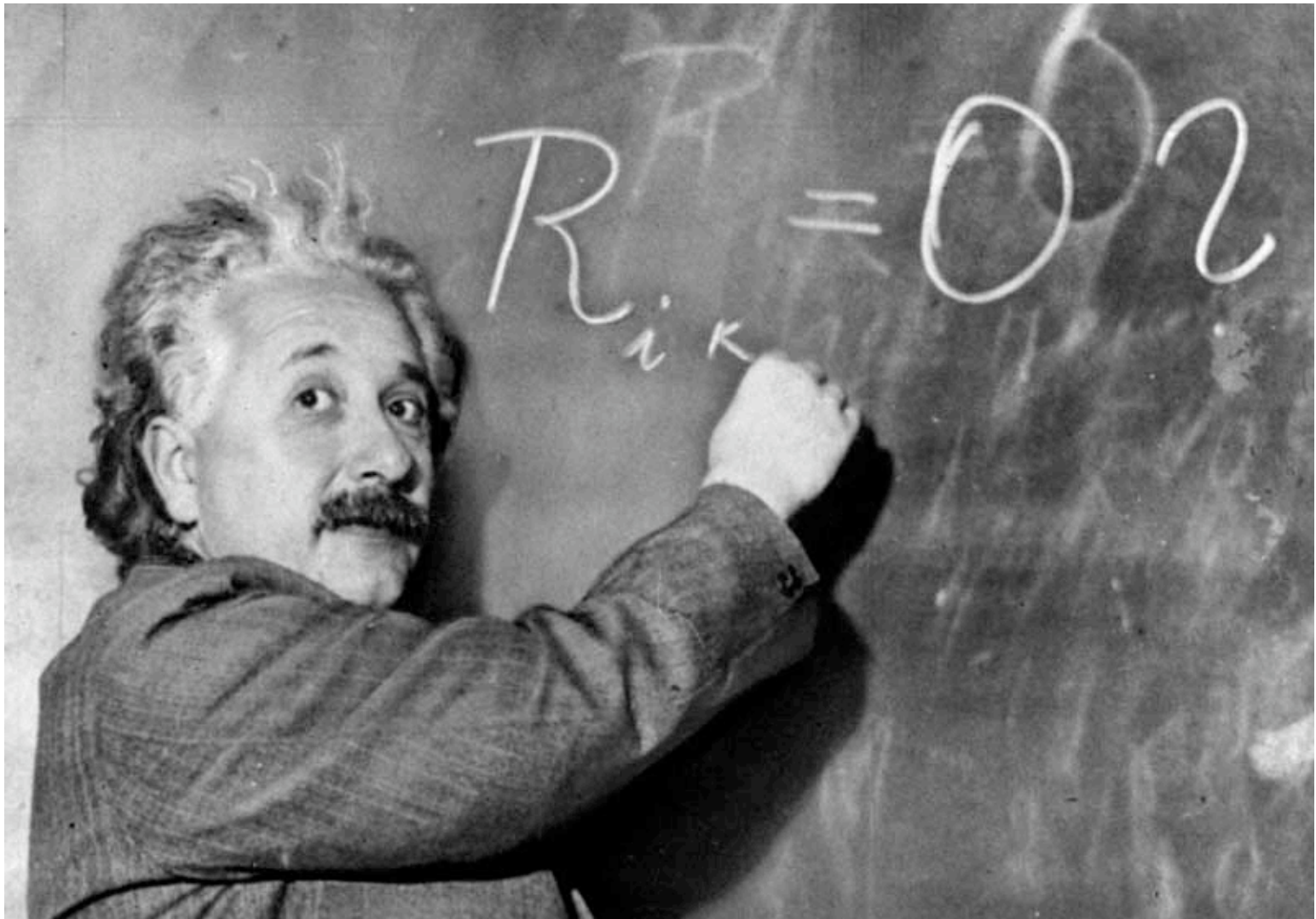
“... the greatest blunder of my life!”

— A. Einstein, to G. Gamow

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Has same sign as the *scalar curvature*

$$s = r_j^j = \mathcal{R}^{ij}_{ij}.$$

Dimension Four is Exceptional

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When $n = 4$, Einstein metrics satisfy a remarkable conformally-invariant condition.

On Riemannian n -manifold (M, g) ,

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$$\mathcal{R}^{ab}_{cd} = W^{ab}_{cd} + \frac{4}{n-2} \overset{\circ}{r} \begin{bmatrix} a & b \\ c & d \end{bmatrix} + \frac{2}{n(n-1)} \mathfrak{s} \delta \begin{bmatrix} a & b \\ c & d \end{bmatrix}$$

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W^a_{bcd} unchanged if $g \rightsquigarrow \hat{g} = u^2 g$.

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Proposition. Assume $n \geq 4$. Then

(M^n, g) locally conformally flat $\iff W \equiv 0$.

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Measures deviation $[g]$ from conformal flatness.

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Of course, conformally Einstein good enough!

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But when $n \neq 4$, Einstein $\not\Rightarrow$ critical point of \mathcal{W} !

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When $n = 4$, conf. Einstein \implies critical for \mathcal{W} .

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$$\mathcal{R} = \left(\begin{array}{c|c} W_+ + \frac{s}{12} & \mathring{r} \\ \hline \mathring{r} & W_- + \frac{s}{12} \end{array} \right)$$

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	Λ^{+*}	Λ^{-*}
Λ^+	$W_+ + \frac{s}{12}$	\mathring{r}
Λ^-	\mathring{r}	$W_- + \frac{s}{12}$

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$$\tau(M) = \frac{1}{12\pi^2} \int_M \left(|W_+|^2 - |W_-|^2 \right) d\mu$$

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Hence

$$\mathcal{W}([g]) = -12\pi^2 \tau(M) + 2 \int_M |W_+|^2 d\mu_g$$

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A case of special interest:

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(M^4, g, J) Kähler.

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$$|W_+|^2 = \frac{s^2}{24}$$

Restriction of \mathcal{W}_+ to Kähler metrics?

On Kähler metrics,

$$\int |W_+|^2 d\mu = \int \frac{s^2}{24} d\mu$$

so any critical point of restriction must be **extremal** in sense of Calabi.

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Andrzej Derdziński : For Kähler metrics g ,

$$B = \frac{1}{12} \left[2s\dot{r} + \text{Hess}_0(s) + 3J^* \text{Hess}_0(s) \right]$$

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- $g_t = g + tB$ is Kähler metric for small t .

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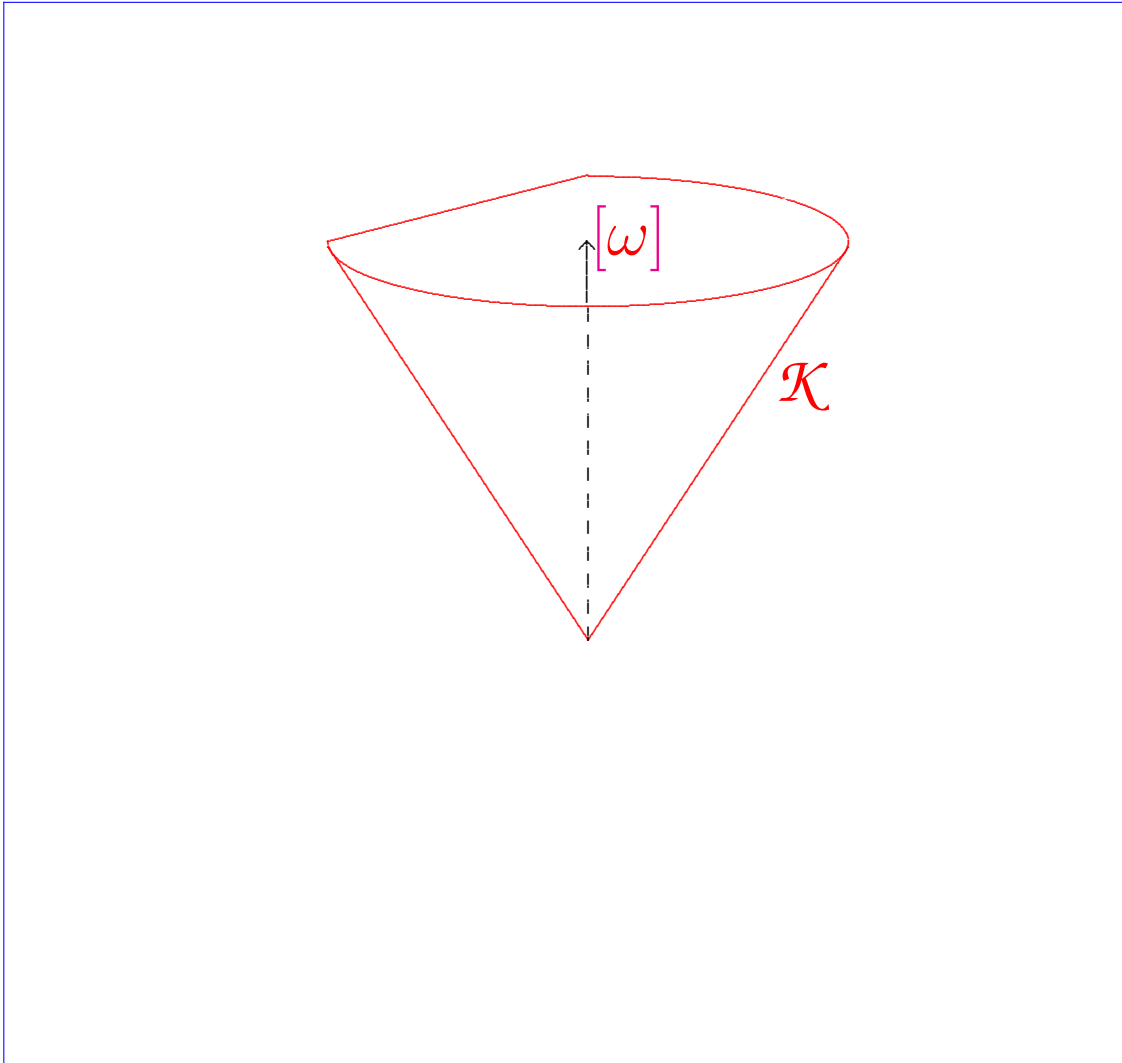
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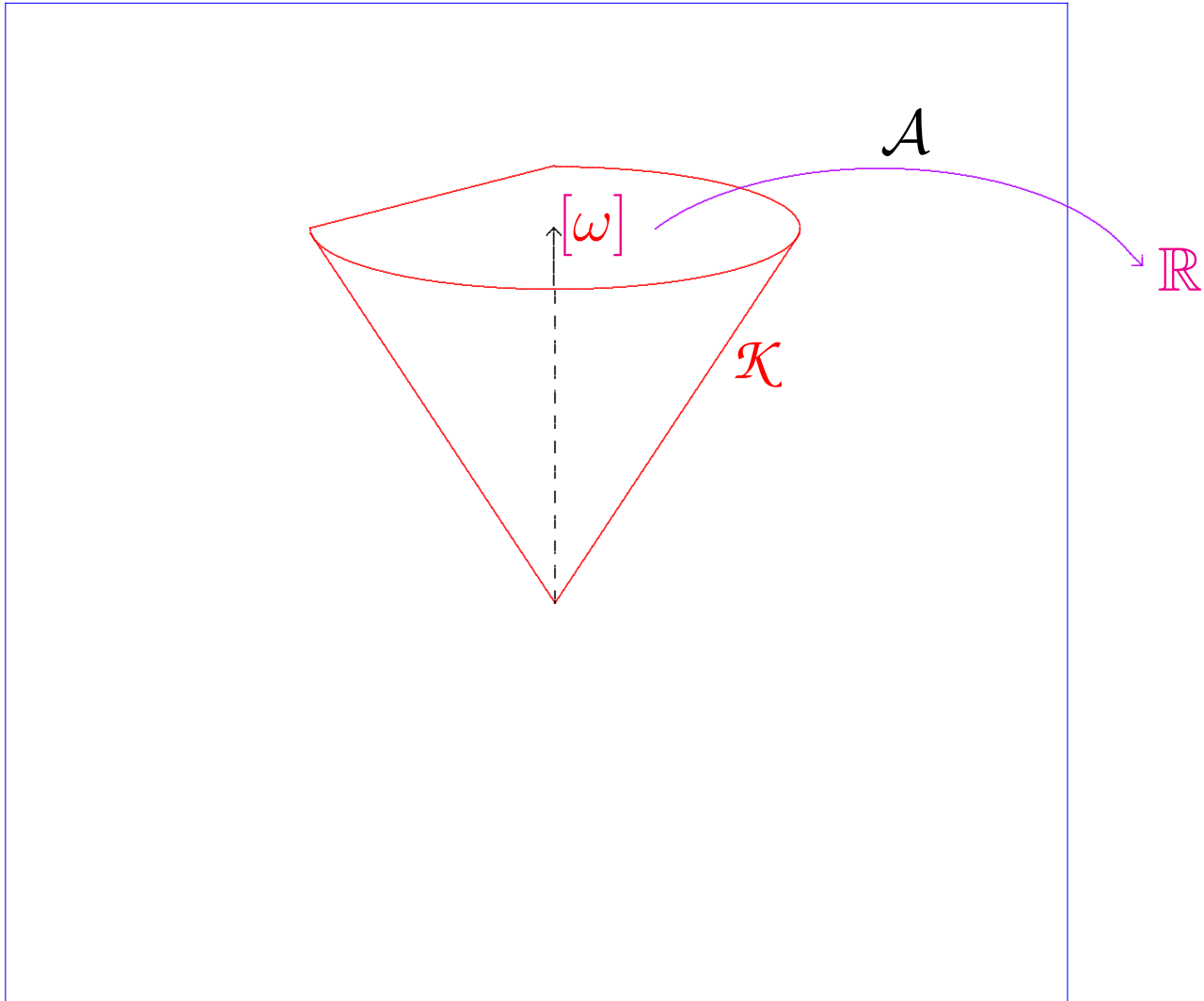
For any extremal Kähler (M^4, g, J) ,

$$\begin{aligned} \frac{1}{32\pi^2} \int s^2 d\mu_g &= \frac{(c_1 \cdot [\omega])^2}{[\omega]^2} + \frac{1}{32\pi^2} \|\mathcal{F}_{[\omega]}\|^2 \\ &=: \mathcal{A}([\omega]) \end{aligned}$$

where \mathcal{F} is Futaki invariant.



$$\mathcal{K} \subset H^{1,1}(M, \mathbb{R}) \subset H^2(M, \mathbb{R})$$



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Proposition. *If g is a Kähler metric on a compact complex surface (M^4, J) , with Kähler class $[\omega]$,*

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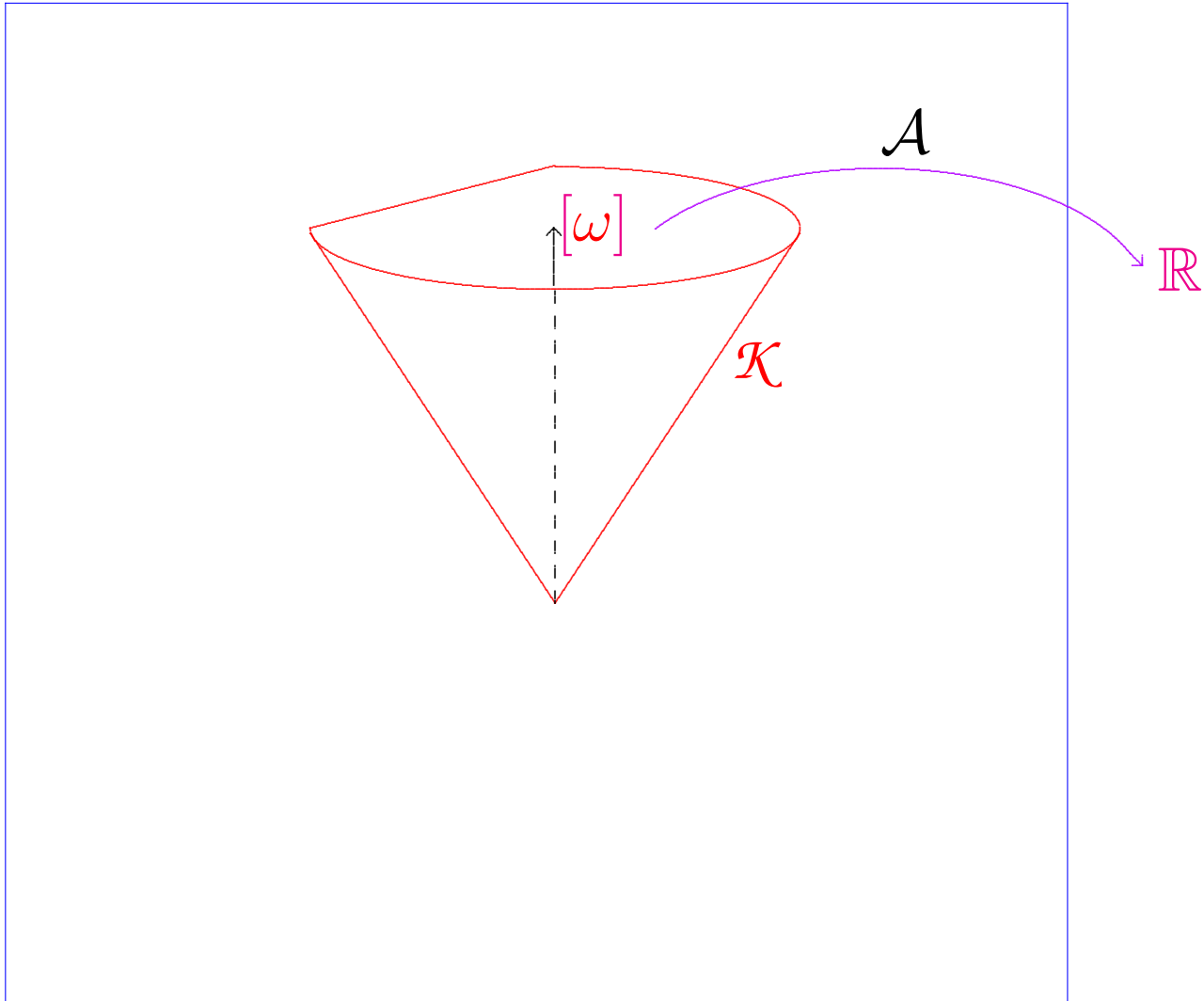
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Restriction of \mathcal{W}_+ to Kähler metrics?

On Kähler metrics,

$$\int |W_+|^2 d\mu = \int \frac{s^2}{24} d\mu$$

so any critical point of restriction must be **extremal** in sense of Calabi.

Andrzej Derdziński : For Kähler metrics g ,

$$B = \frac{1}{12} \left[2s\dot{r} + \text{Hess}_0(s) + 3J^* \text{Hess}_0(s) \right]$$

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Global implications?

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Calabi-Eckmann: False in higher dimensions!

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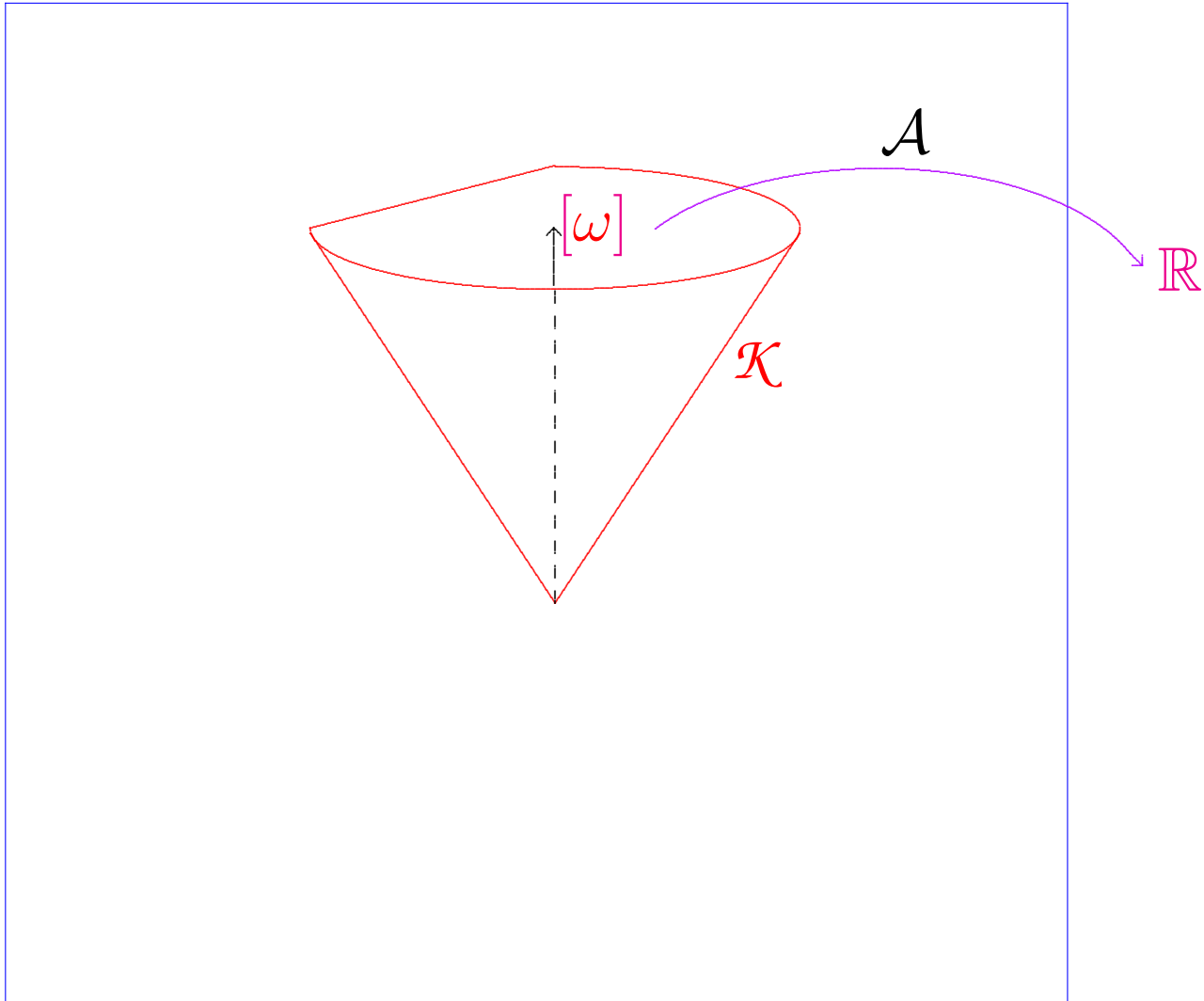
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Implies uniqueness result for h on (M^4, J) .

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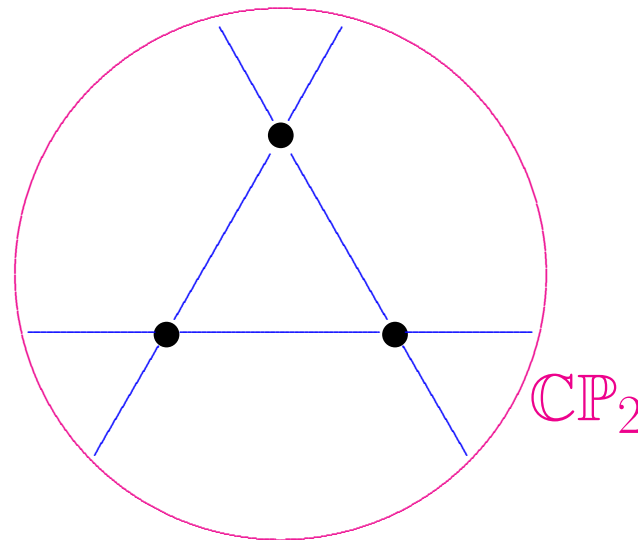
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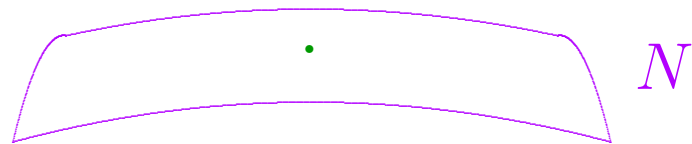
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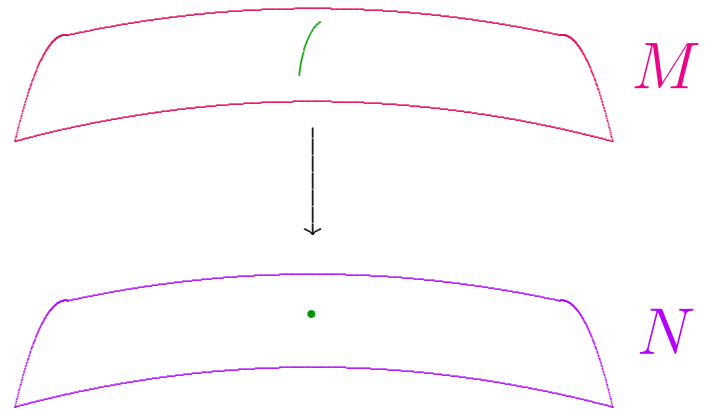
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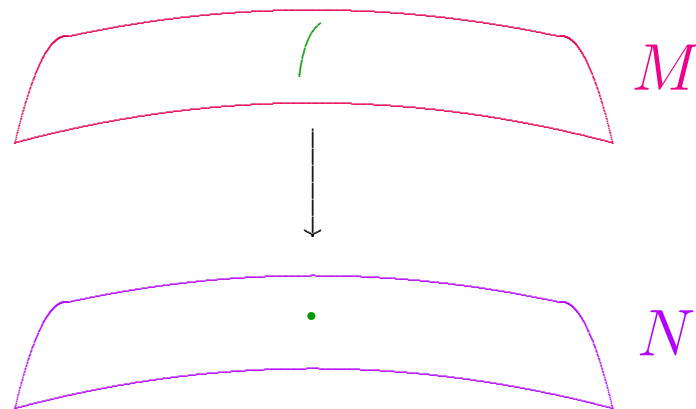
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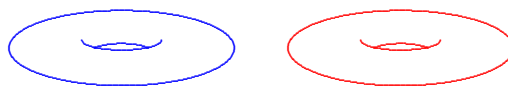
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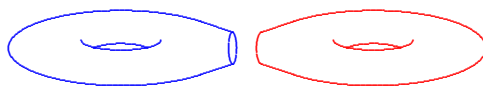
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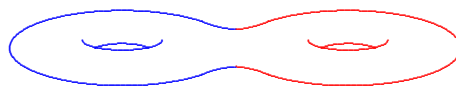
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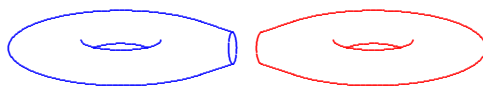
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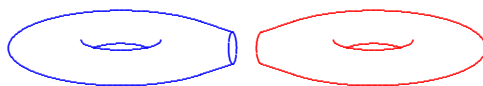
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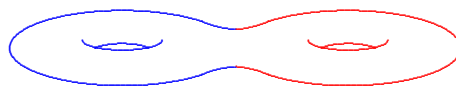
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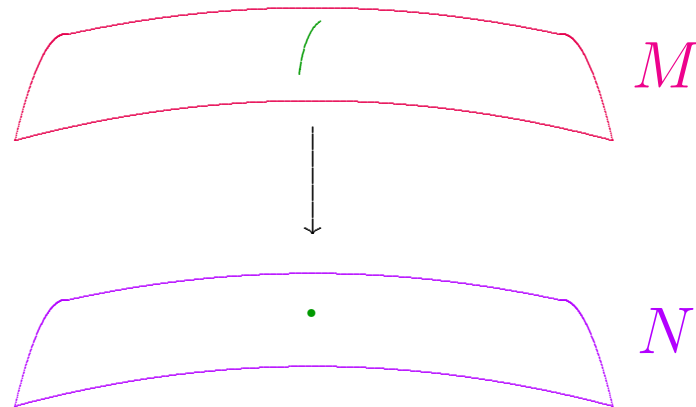
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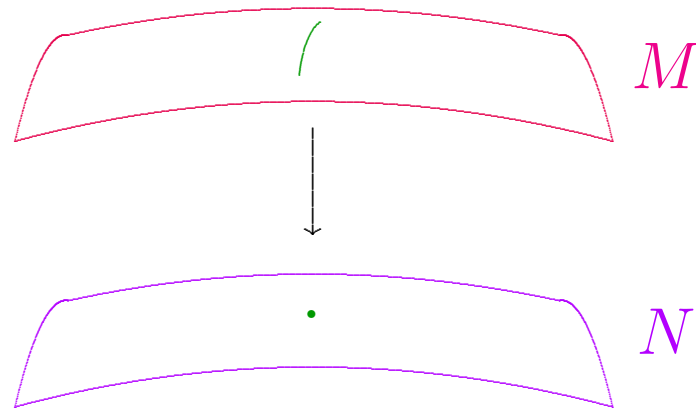


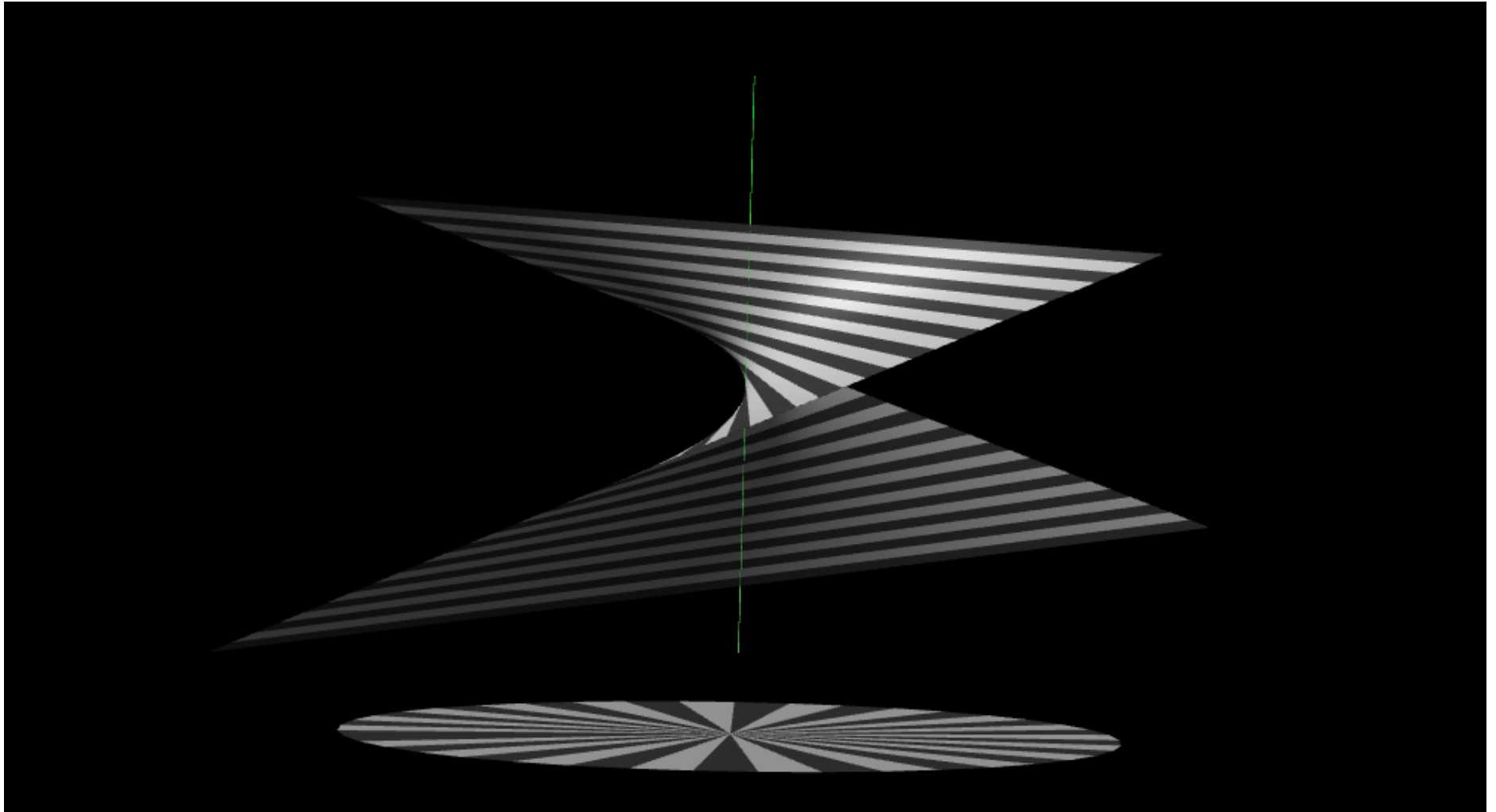
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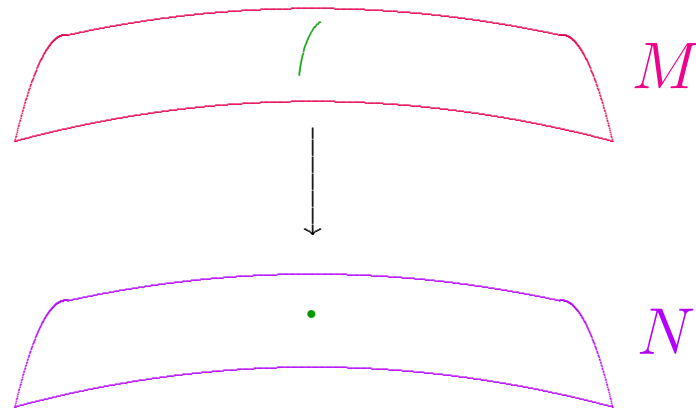


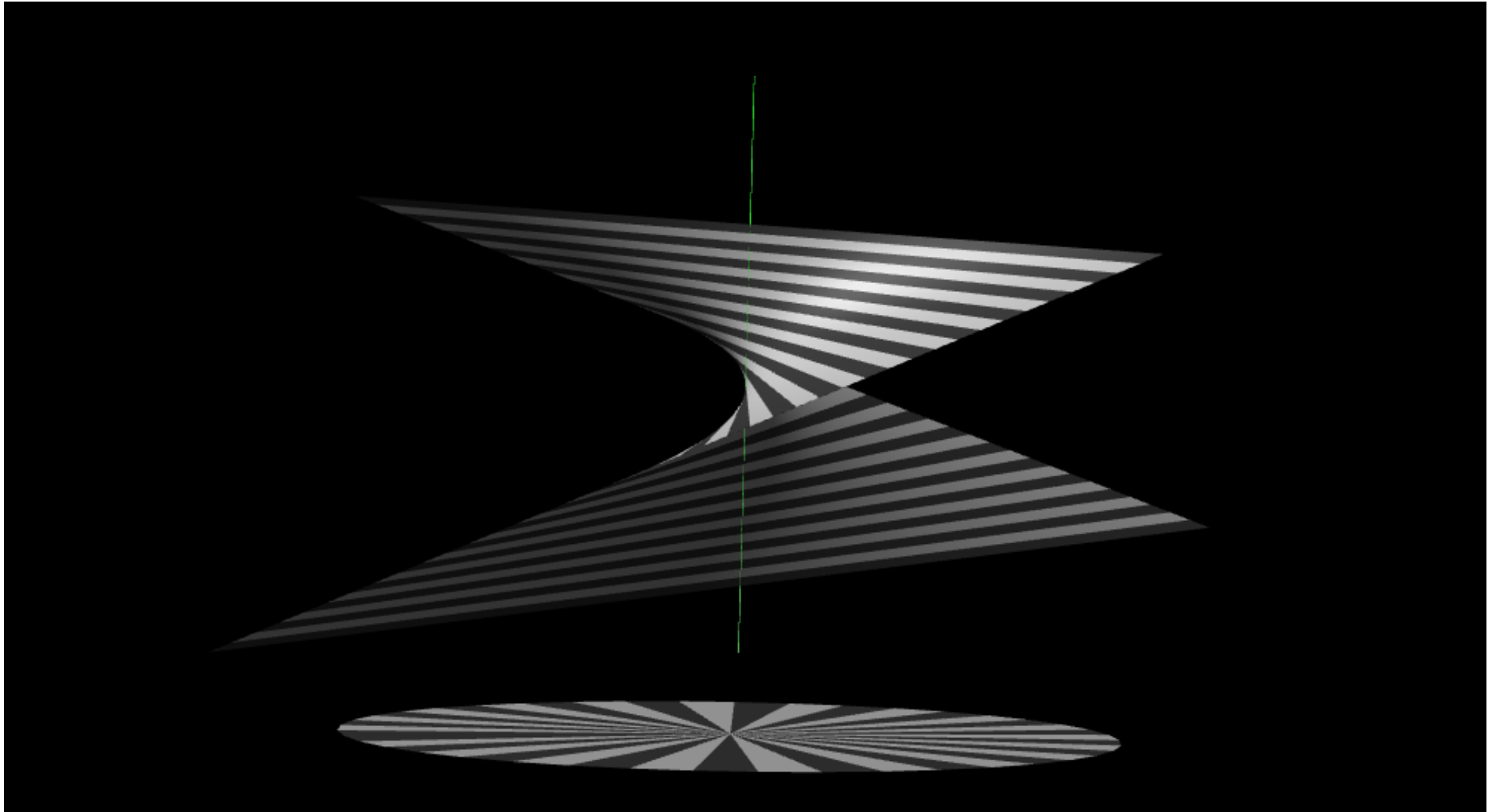
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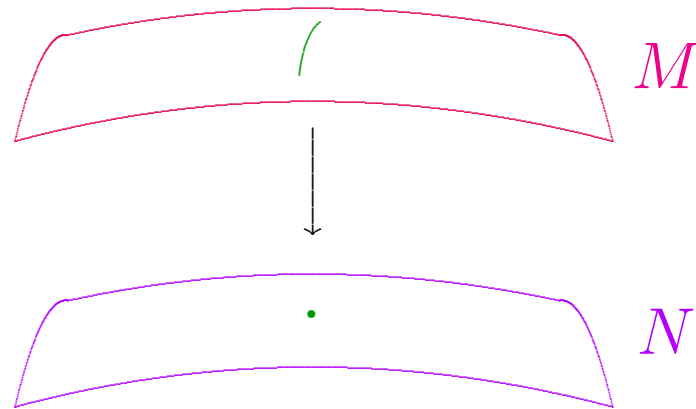


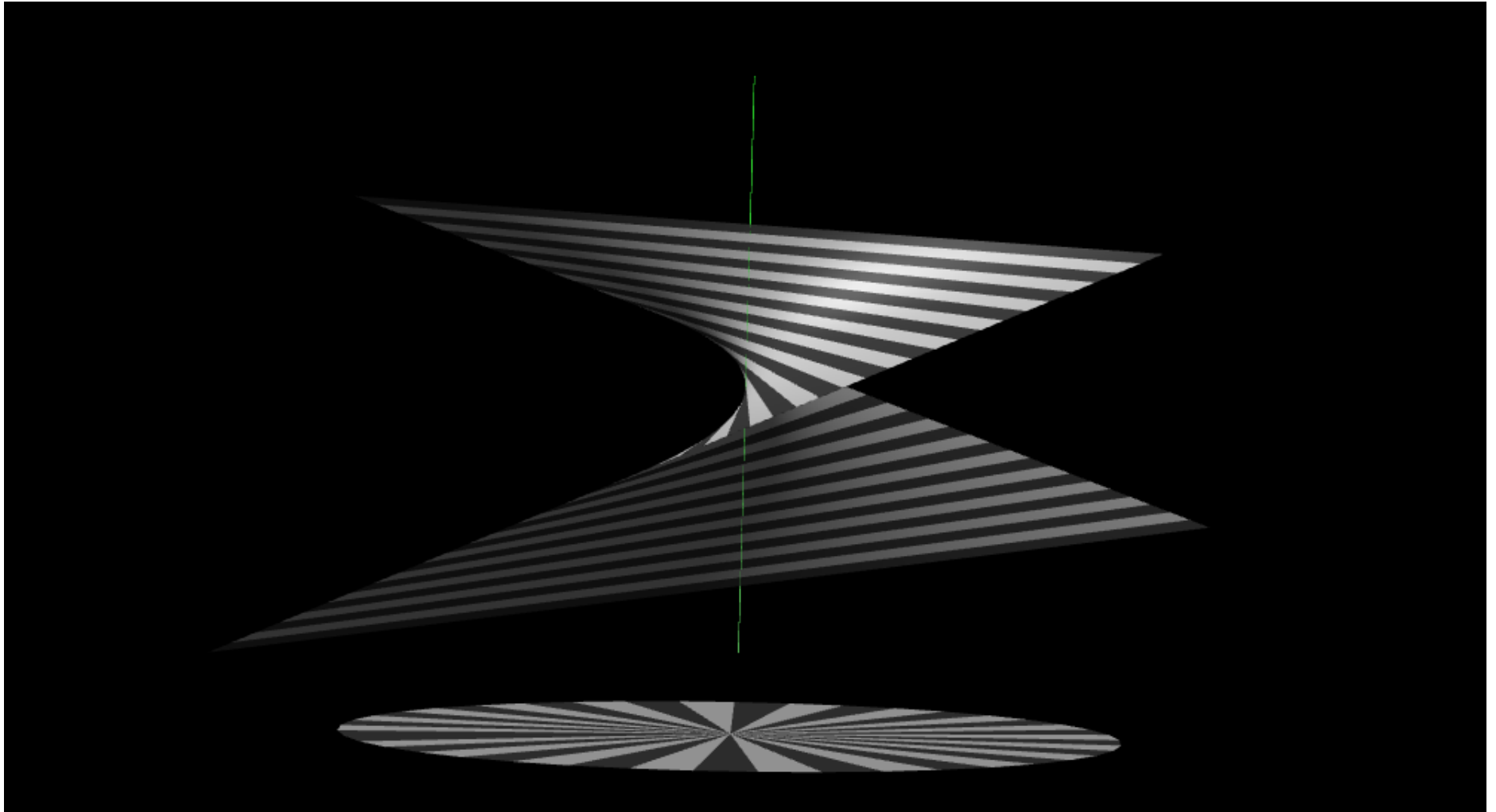
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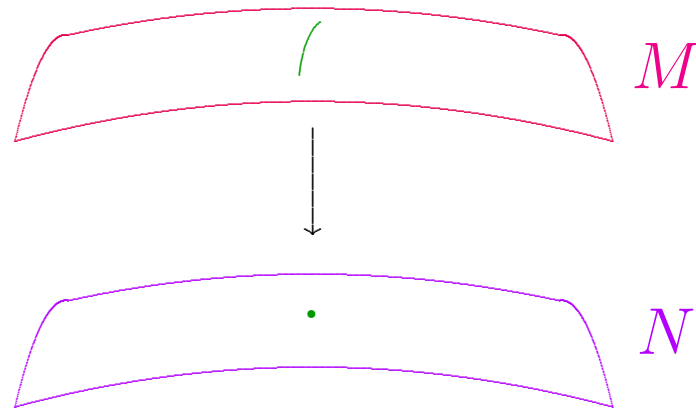


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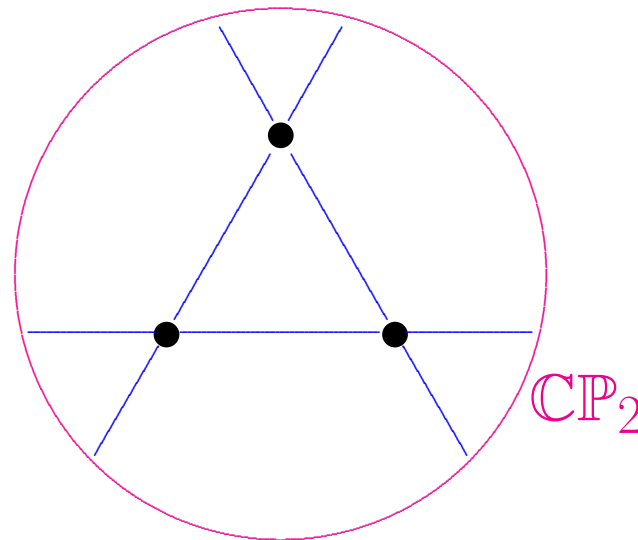


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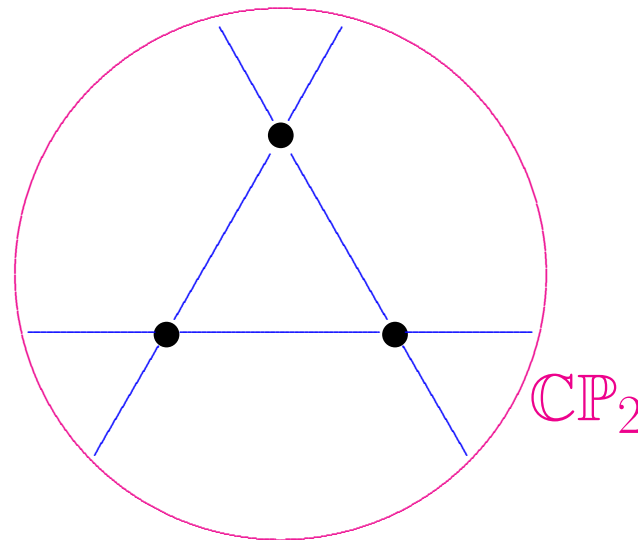
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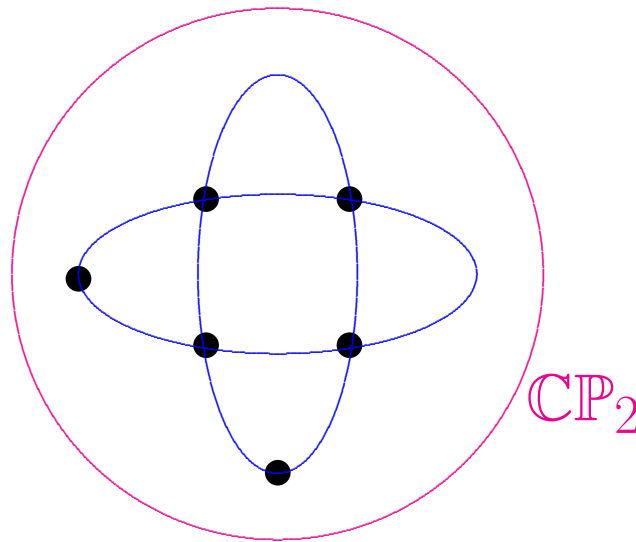


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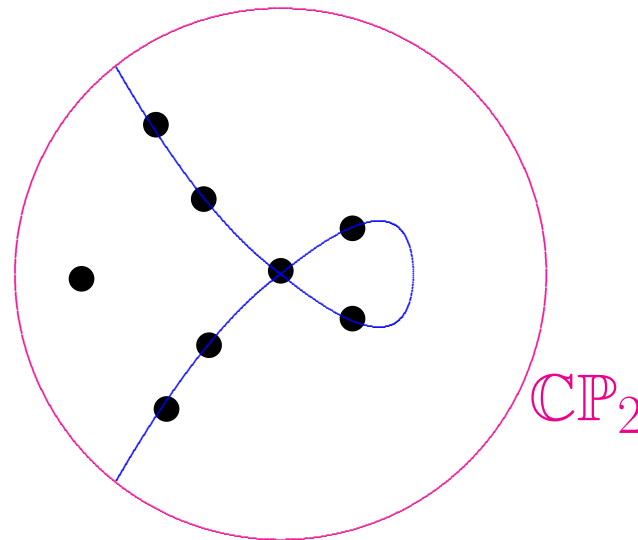


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No 3 on a line, no 6 on conic, no 8 on nodal cubic.

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Theorem. *Each del Pezzo (M^4, J) admits a J -compatible*

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One reason this seems satisfying...

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Exactly one connected component of moduli space!

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Claim: (M, h) compact Einstein $\implies J$ integrable.

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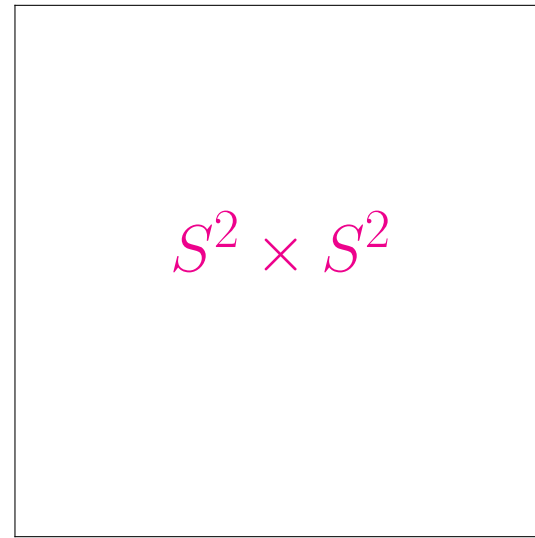
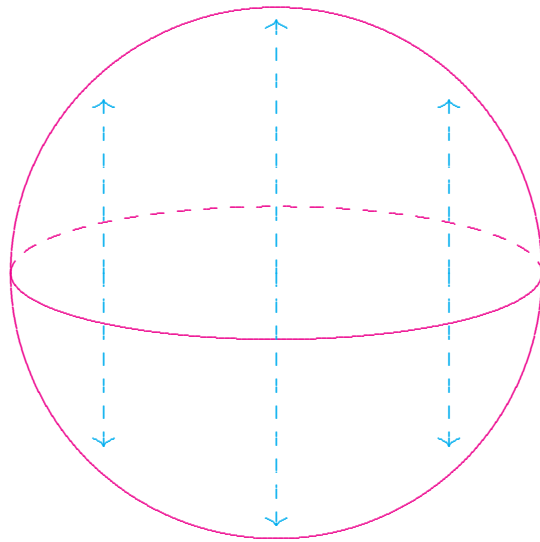
Simply connected hypothesis is essential!

Theorem B. *Let M be smooth compact oriented 4-manifold with $\pi_1 \neq 0$.*

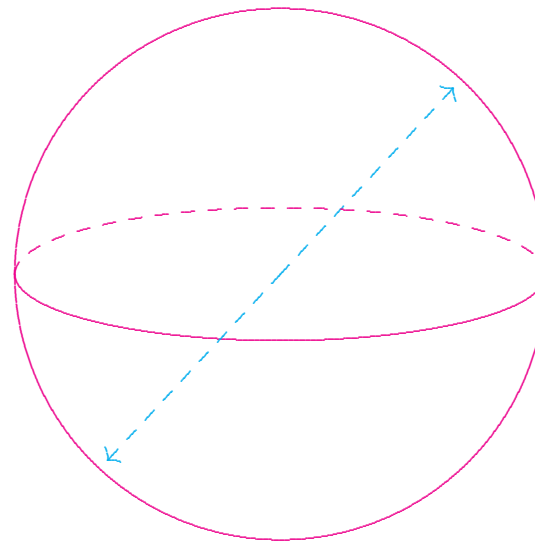
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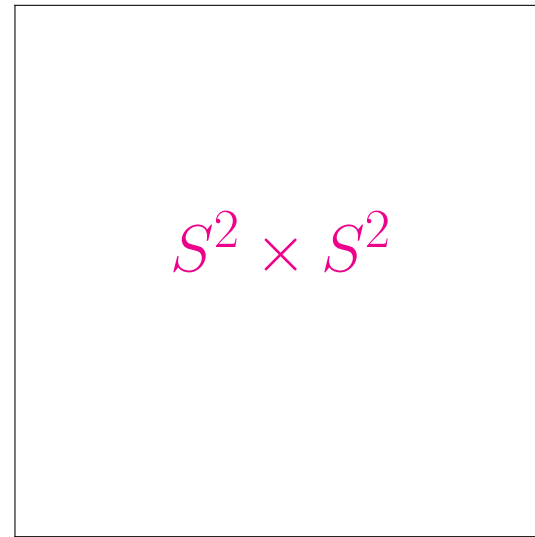
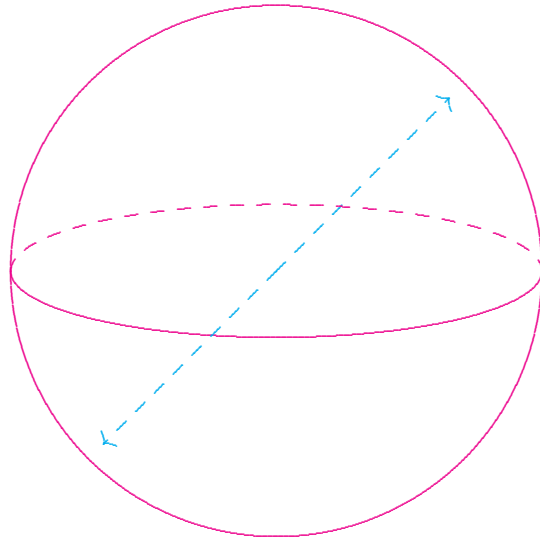


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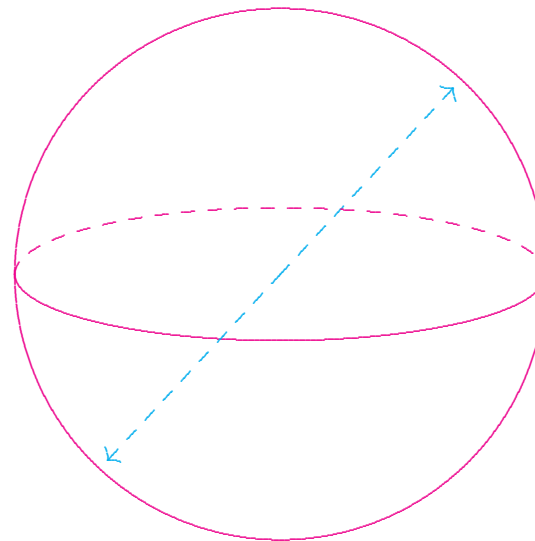
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Non-spin 4-manifold

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Why is $\mathcal{E}_{\det}(M) \subset \mathcal{E}(M)$ open and closed?

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Theorem D. *Let (M, h) be a compact oriented Einstein 4-manifold. If*

$$\det(W^+) > -\frac{5\sqrt{2}}{21\sqrt{21}}|W^+|^3$$

everywhere on M , then actually $\det(W^+) > 0$. Consequently, all the results described remain true if we merely impose this ostensibly weaker hypothesis.

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$$0 = \int_M \left[\langle W^+, \nabla^* \nabla (\omega \otimes \omega) \rangle + \frac{s}{2} W^+ (\omega, \omega) - 6 |W^+ (\omega)|^2 + 2 |W^+|^2 |\omega|^2 \right] f d\mu$$

Let $\alpha \geq \beta \geq \gamma$ be eigenvalues of W^+ :

$$W^+ = \begin{bmatrix} \alpha & & \\ & \beta & \\ & & \gamma \end{bmatrix}$$

$$\alpha + \beta + \gamma = 0$$

$$\alpha > 0, \quad \gamma < 0, \quad \text{if } W^+ \neq 0$$

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$$f = \alpha_h^{-1/3}, \quad g = f^{-2}h = \alpha_h^{2/3}h.$$

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Now choose $\omega \in \Gamma\Lambda^+$ so that

$$W_g^+(\omega) = \alpha \omega, \quad |\omega|_g \equiv \sqrt{2},$$

after at worst passing to double cover $\hat{M} \rightarrow M$.

$$0 = \int_{\hat{M}} \left[\langle W^+, \nabla^* \nabla (\omega \otimes \omega) \rangle + \frac{s}{2} W^+(\omega, \omega) - 6 |W^+(\omega)|^2 + 2 |W^+|^2 |\omega|^2 \right] f \, d\mu$$

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$$0 = \int_M \left[-2W^+(\nabla_e \omega, \nabla^e \omega) - 2W^+(\omega, \nabla^e \nabla_e \omega) \right. \\ \left. + \frac{s}{2} W^+(\omega, \omega) - 6|W^+(\omega)|^2 + 2|W^+|^2 |\omega|^2 \right] f \, d\mu$$

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because

$$W_g^+(\omega) = \alpha \omega$$

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because

$$|W_g^+|^2 \geq \frac{3}{2} \alpha^2$$

$$0 \geq \int_M \left[-2W^+(\nabla_e \omega, \nabla^e \omega) + 2\alpha \langle \omega, \nabla^* \nabla \omega \rangle \right. \\ \left. + \frac{s}{2} \alpha |\omega|^2 - 3\alpha^2 |\omega|^2 \right] f \, d\mu$$

$$|\omega|_g^2 = 2 \implies (\nabla_e \omega) \perp \omega$$

$$0 \geq \int_M \left[-2W^+(\nabla_e \omega, \nabla^e \omega) + 2\alpha \langle \omega, \nabla^* \nabla \omega \rangle \right. \\ \left. + \frac{s}{2} \alpha |\omega|^2 - 3\alpha^2 |\omega|^2 \right] f \, d\mu$$

$$|\omega|_g^2 = 2 \implies (\nabla_e \omega) \perp \omega$$

$$\det(W^+) > 0 \implies W^+ \sim \begin{bmatrix} + & & \\ & - & \\ & & - \end{bmatrix}$$

$$0 \geq \int_M \left[-2W^+(\nabla_e \omega, \nabla^e \omega) + 2\alpha \langle \omega, \nabla^* \nabla \omega \rangle \right. \\ \left. + \frac{s}{2} \alpha |\omega|^2 - 3\alpha^2 |\omega|^2 \right] f \, d\mu$$

$$|\omega|_g^2 = 2 \implies (\nabla_e \omega) \perp \omega$$

$$\det(W^+) > 0 \implies W^+(\nabla_e \omega, \nabla^e \omega) \leq 0$$

$$0 \geq \int_M \left[\begin{aligned} &2\alpha \langle \omega, \nabla^* \nabla \omega \rangle \\ &+ \frac{s}{2} \alpha |\omega|^2 - 3\alpha^2 |\omega|^2 \end{aligned} \right] f \, d\mu$$

$$|\omega|_g^2 = 2 \implies (\nabla_e \omega) \perp \omega$$

$$\det(W^+) > 0 \implies -W^+(\nabla_e \omega, \nabla^e \omega) \geq 0$$

$$0 \geq \int_M \left[2\alpha \langle \omega, \nabla^* \nabla \omega \rangle + \frac{s}{2} \alpha |\omega|^2 - 3\alpha^2 |\omega|^2 \right] f \, d\mu$$

$$0 \geq \int_M \left[2\langle \omega, \nabla^* \nabla \omega \rangle + \frac{s}{2} |\omega|^2 - 3\alpha |\omega|^2 \right] (\alpha f) d\mu$$

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But

$$\alpha f \equiv 1$$

$$0 \geq \int_M \left[2\langle \omega, \nabla^* \nabla \omega \rangle + \frac{s}{2} |\omega|^2 - 3|\omega|^2 \alpha \right] d\mu$$

$$0 \geq \int_M \left[2\langle \omega, \nabla^* \nabla \omega \rangle - 3W^+(\omega, \omega) + \frac{s}{2} |\omega|^2 \right] d\mu$$

$$0 \geq \int_M \left[\frac{1}{2} |\nabla \omega|^2 + \frac{3}{2} \langle \omega, \left(\nabla^* \nabla - 2W^+ + \frac{s}{3} \right) \omega \rangle \right] d\mu$$

$$0 \geq \int_M \left[\frac{1}{2} |\nabla \omega|^2 + \frac{3}{2} \langle \omega, (d + d^*)^2 \omega \rangle \right] d\mu$$

Because

$$(d + d^*)^2 = \nabla^* \nabla - 2W^+ + \frac{s}{3}$$

on $\Gamma\Lambda^+$.

$$0 \geq \frac{1}{2} \int_M |\nabla \omega|^2 d\mu + 3 \int_M |d\omega|^2 d\mu$$

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So $\nabla \omega \equiv 0$, and g is Kähler!

$$0 \geq \frac{1}{2} \int_M |\nabla \omega|^2 d\mu + 3 \int_M |d\omega|^2 d\mu$$

So $\nabla \omega \equiv 0$, and g is Kähler!

QED





Thanks for the invitation!



It's such a pleasure to be here!



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