

Einstein Metrics,
Complex Surfaces, &
Symplectic 4-Manifolds

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Colloquium de Mathématiques
Université Paul Sabatier
Toulouse, 10 mai, 2019

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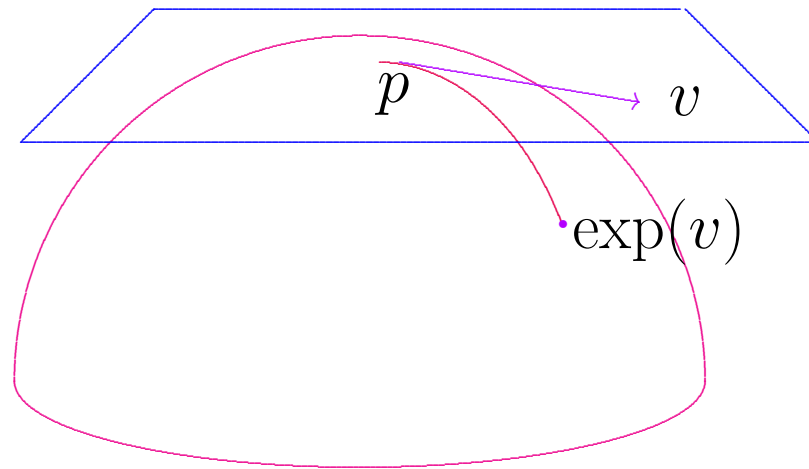
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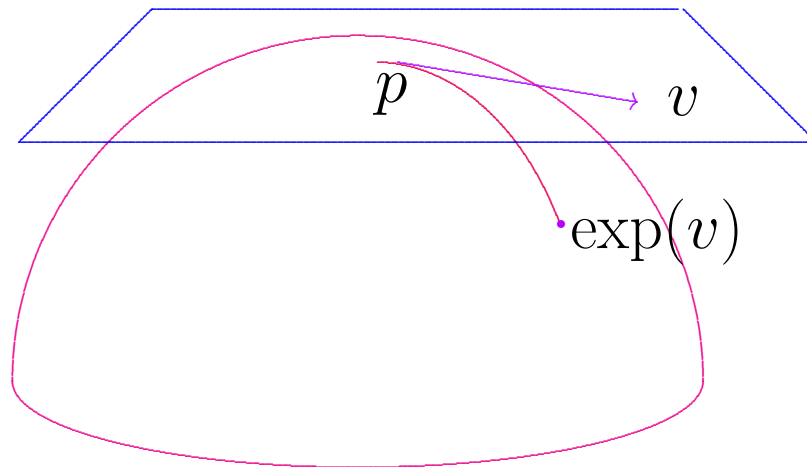
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Now choosing $T_p M \xrightarrow{\cong} \mathbb{R}^n$ via some orthonormal
basis gives us special coordinates on M .

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Components like \mathcal{R}_{1212} are “**sectional curvatures**” ...

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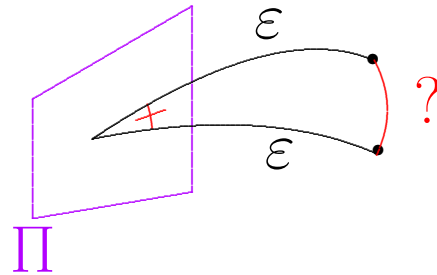
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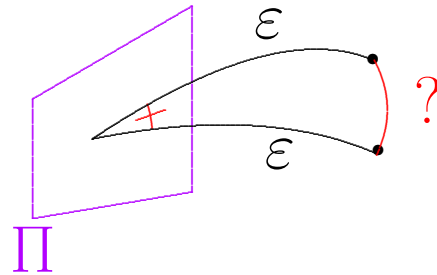


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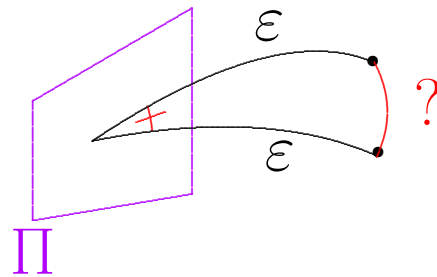
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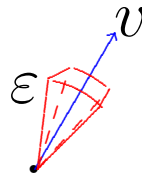
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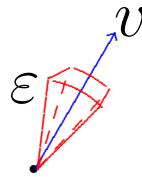


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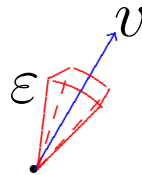
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For example, the unit n -sphere $S^n \subset \mathbb{R}^{n+1}$ has Ricci curvature $\equiv +(n-1)$, but this does not locally characterize the standard metric when $n \geq 4$.

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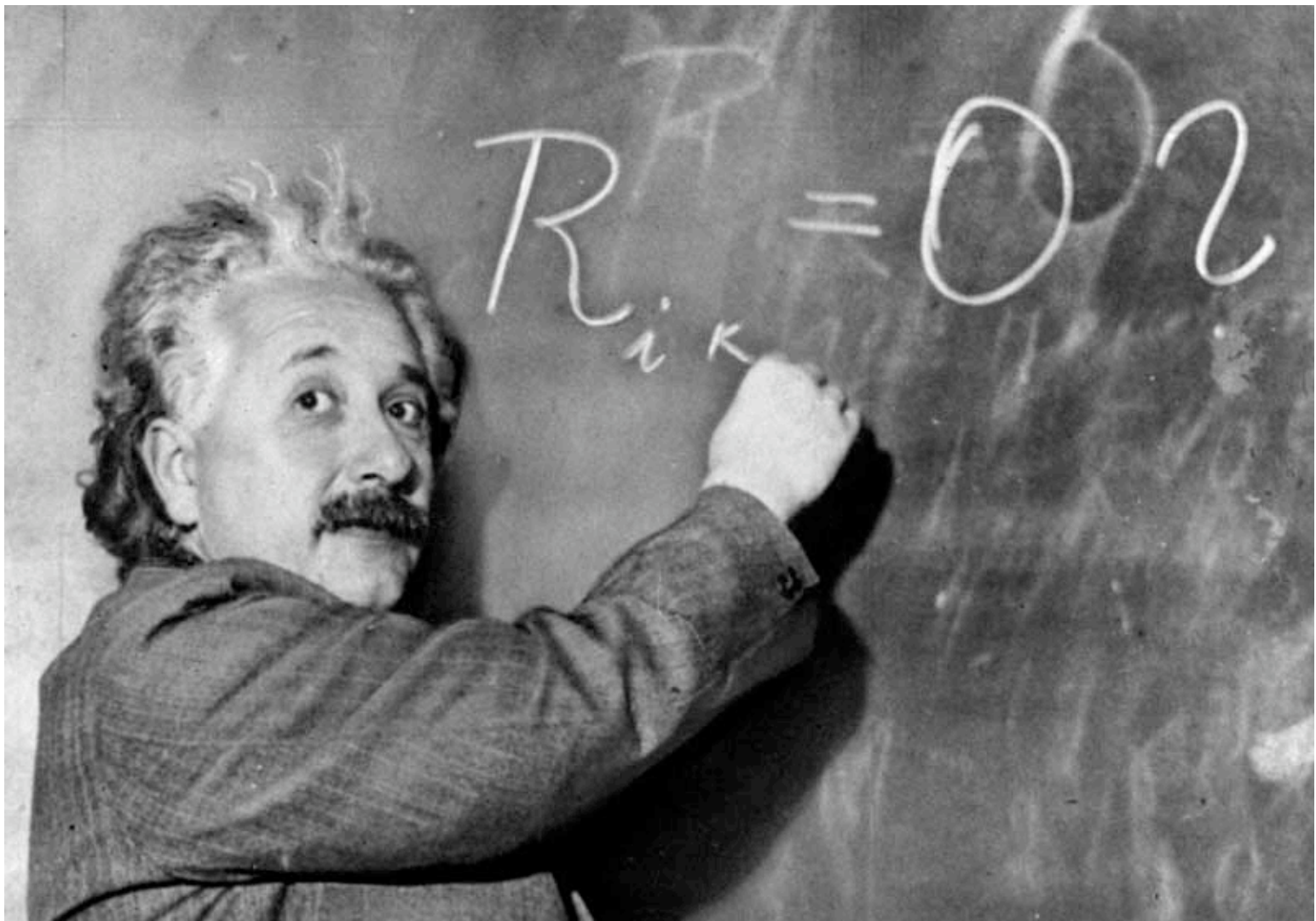
“... the greatest blunder of my life!”

— A. Einstein, to G. Gamow

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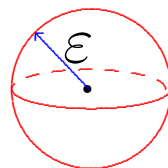
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$$\Delta x^j = 0 \implies r_{jk} = \frac{1}{2} \Delta g_{jk} + \text{lots.}$$

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- When $n \geq 6$, **wide open.** Maybe???

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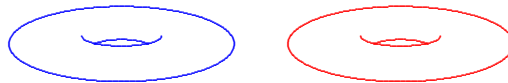
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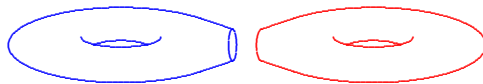
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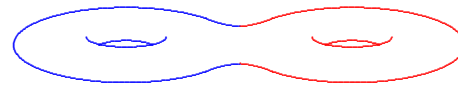
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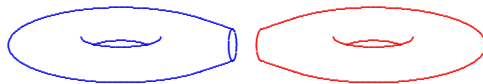
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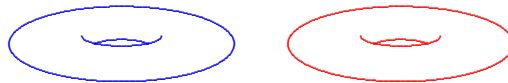
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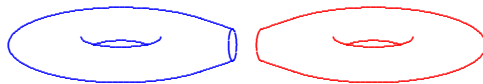
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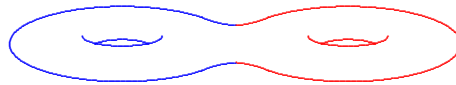
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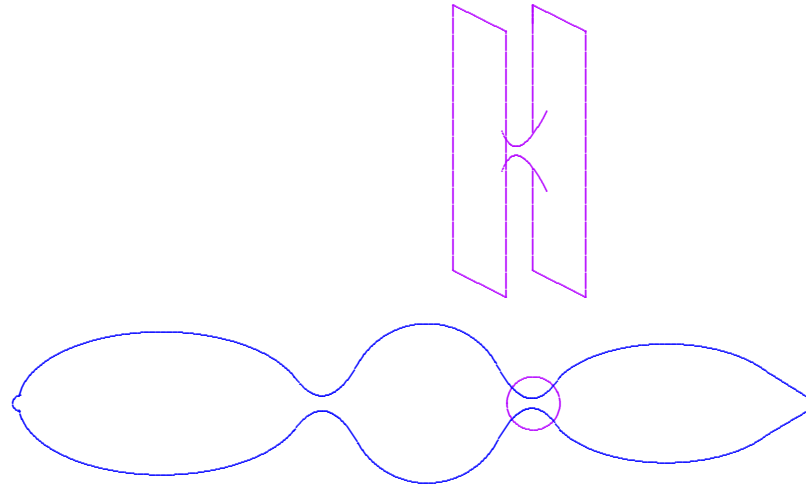
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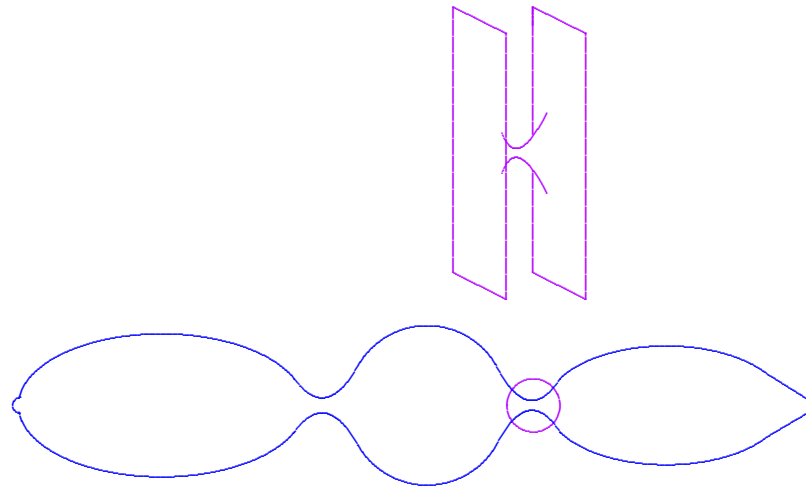
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First step in geometrization:

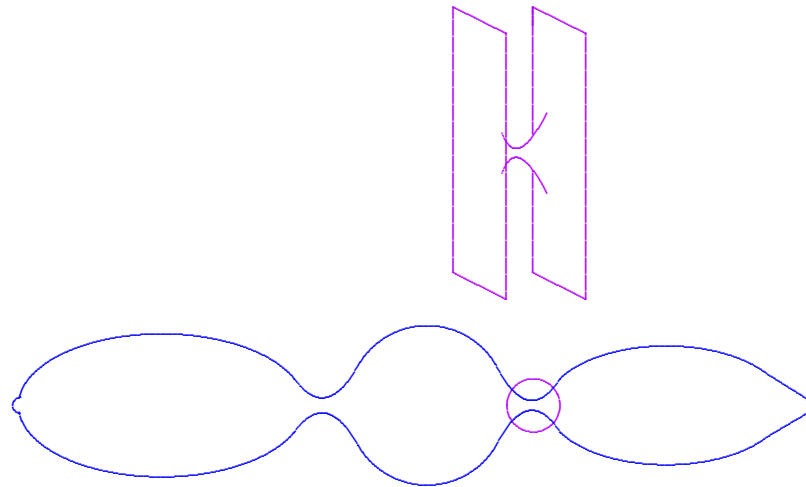
Dimension ≤ 3 :

Einstein's equations are "locally trivial:"

Einstein metrics have constant sectional curvature.

\implies If M^3 carries Einstein metric, $\pi_2(M) = 0$.

\implies Existence obstructed for connect sums $M^3 \# N^3$.



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Prime Decomposition.

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Similar results for most simply connected spin 5-manifolds. (Boyer, Galicki, Kollár, et al.)

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(Terminology to be explained in a moment!)

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K3

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$T^4 =$ Picard torus of curve of genus 2.

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Remove singularities by deforming equation.

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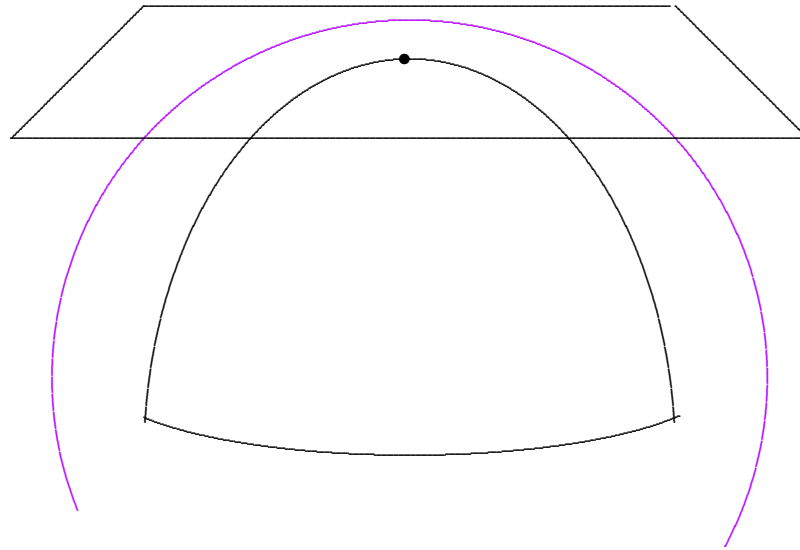
Kähler?

(M^n, g) :

holonomy

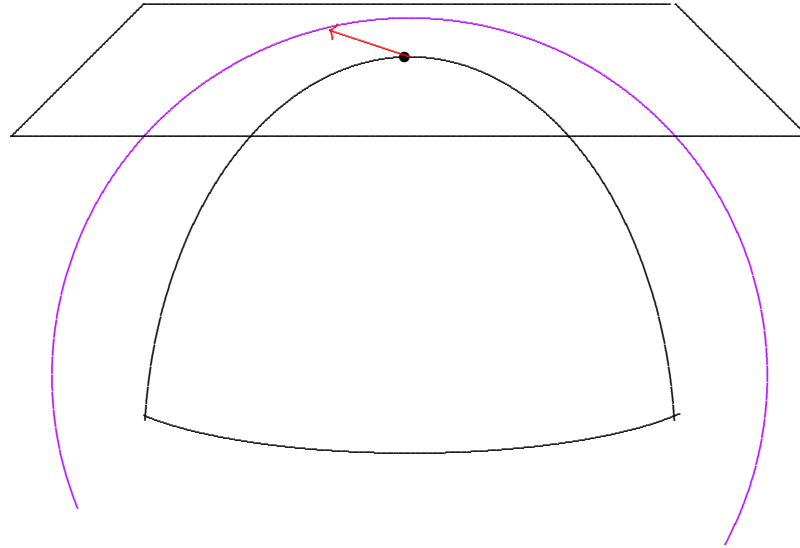
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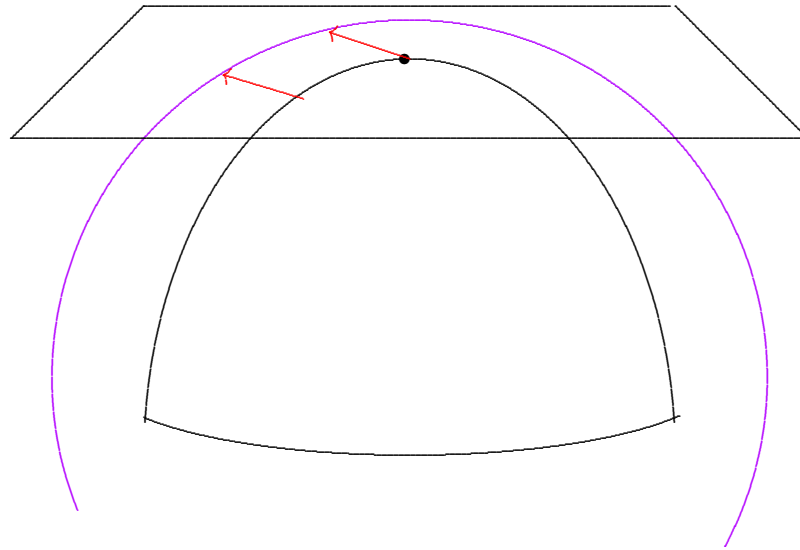
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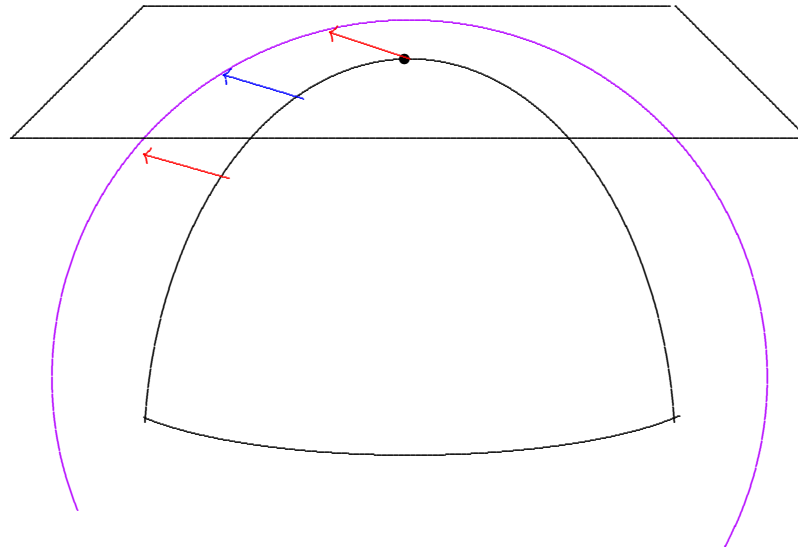
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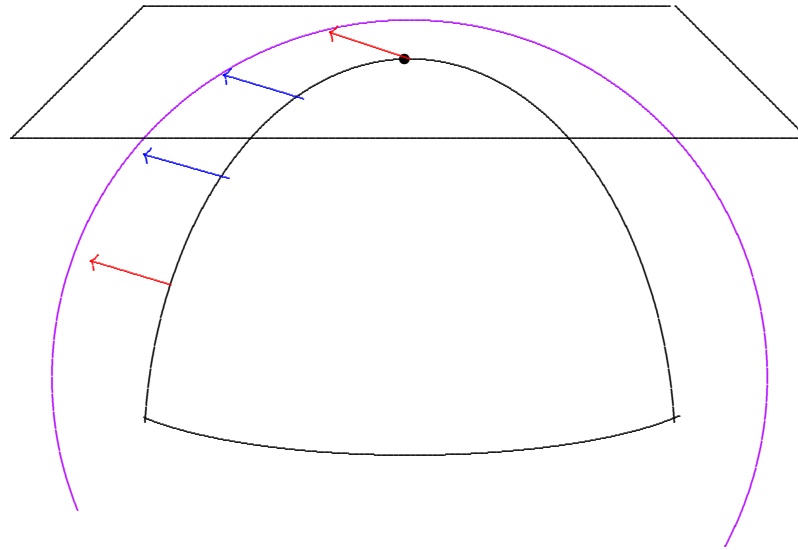
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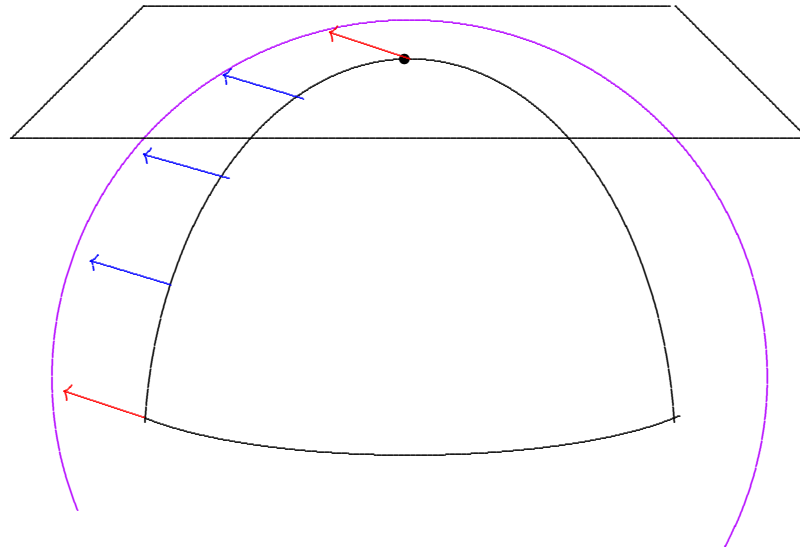
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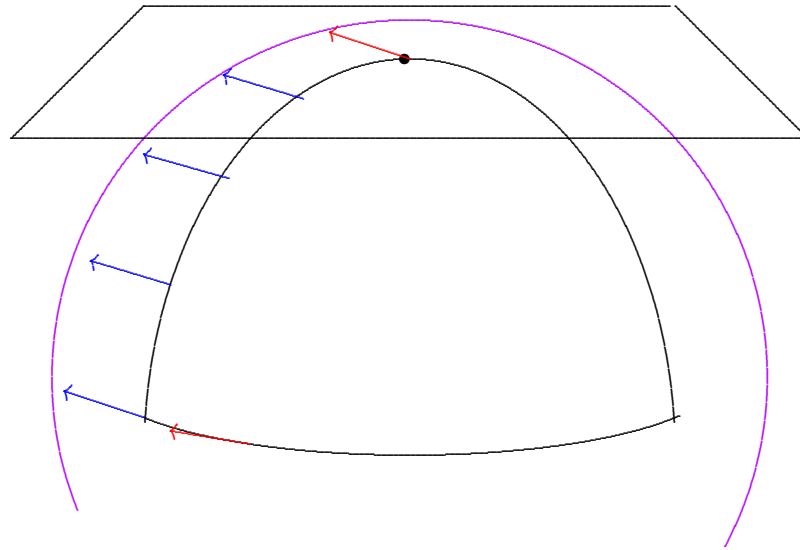
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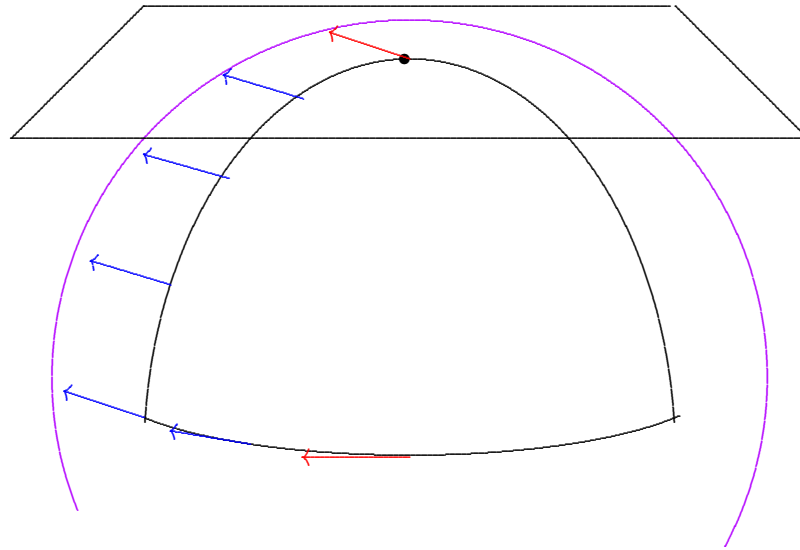
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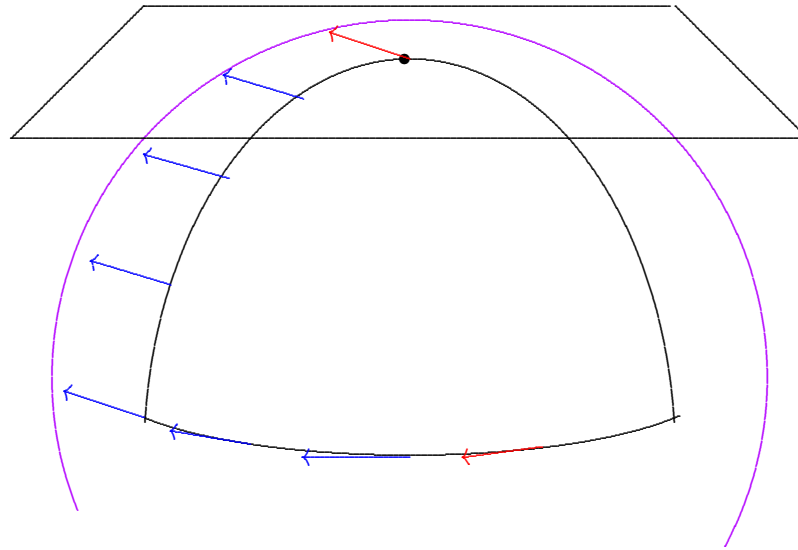
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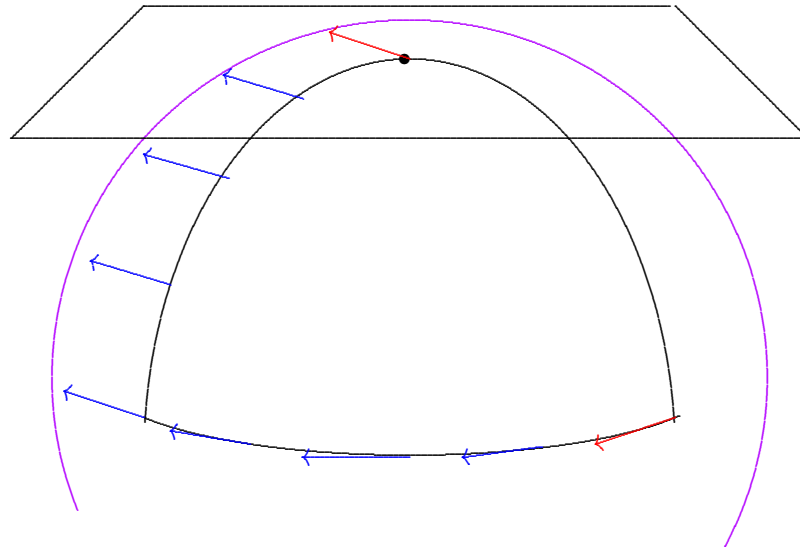
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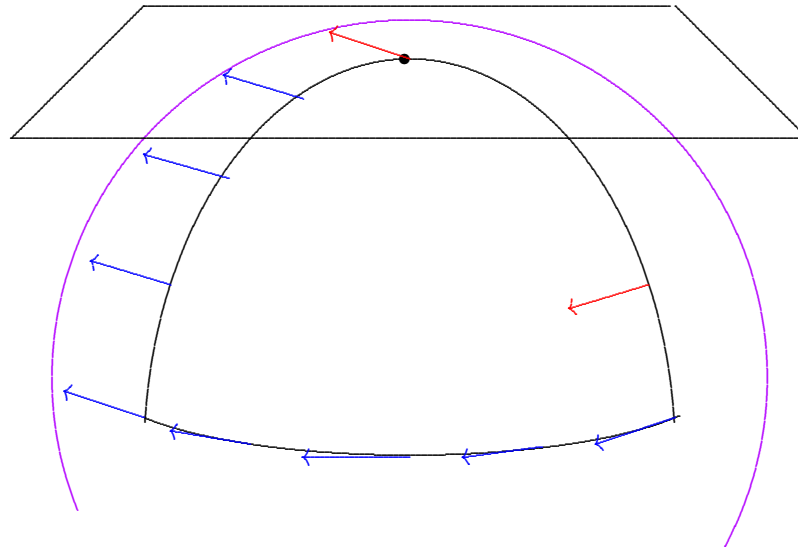
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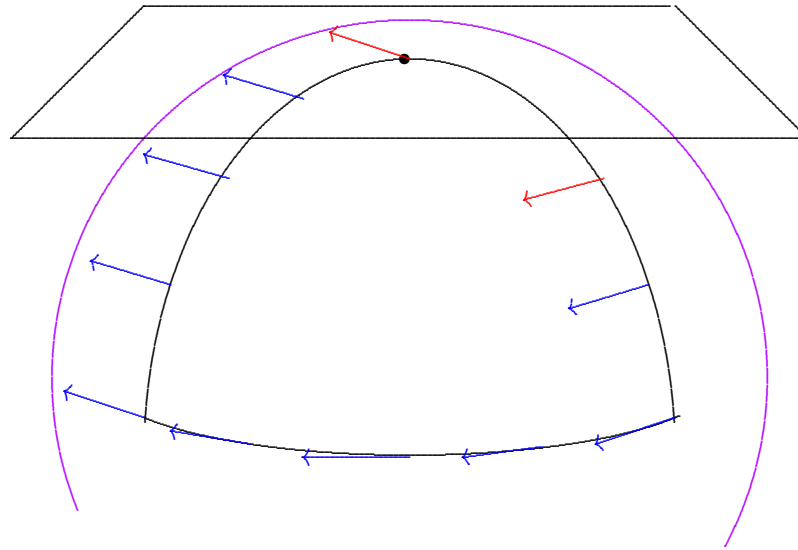
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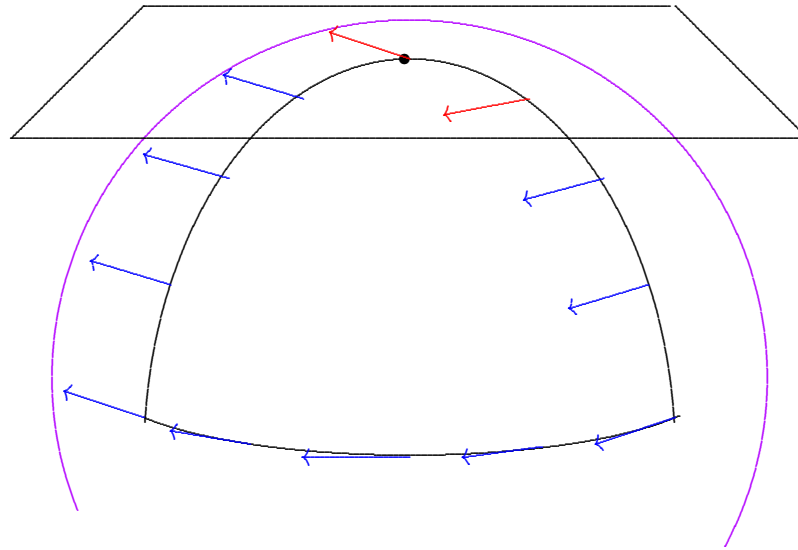
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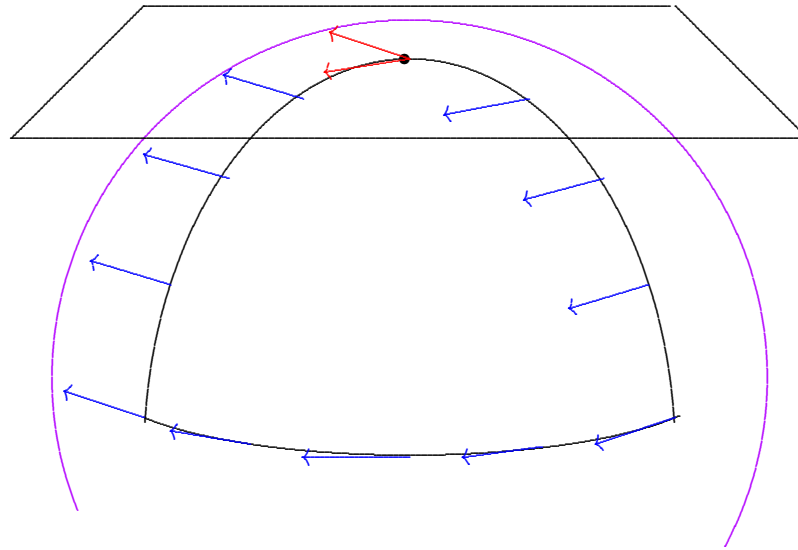
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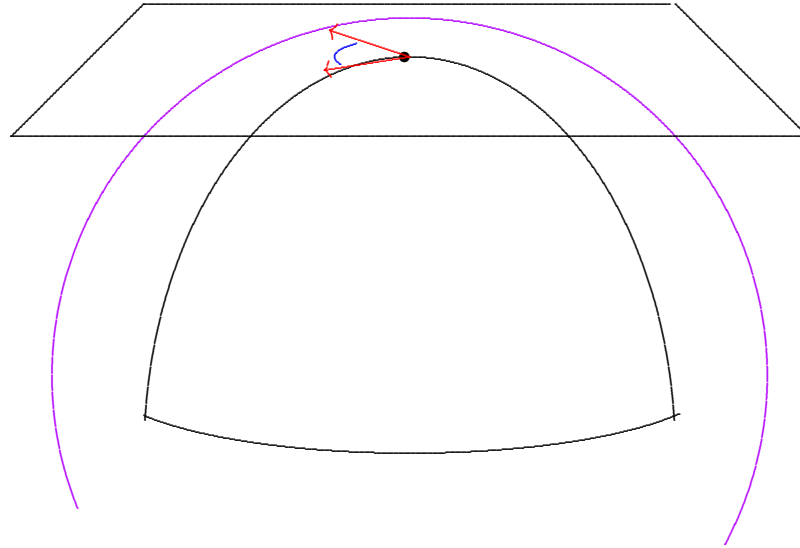
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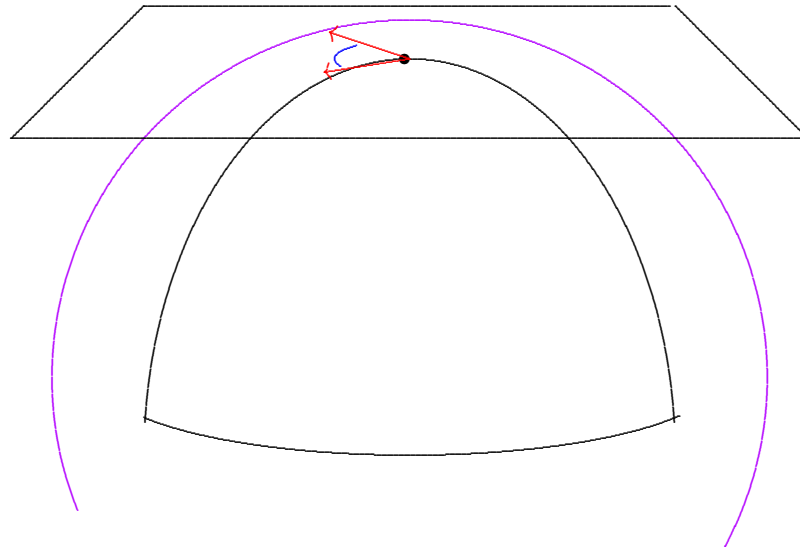
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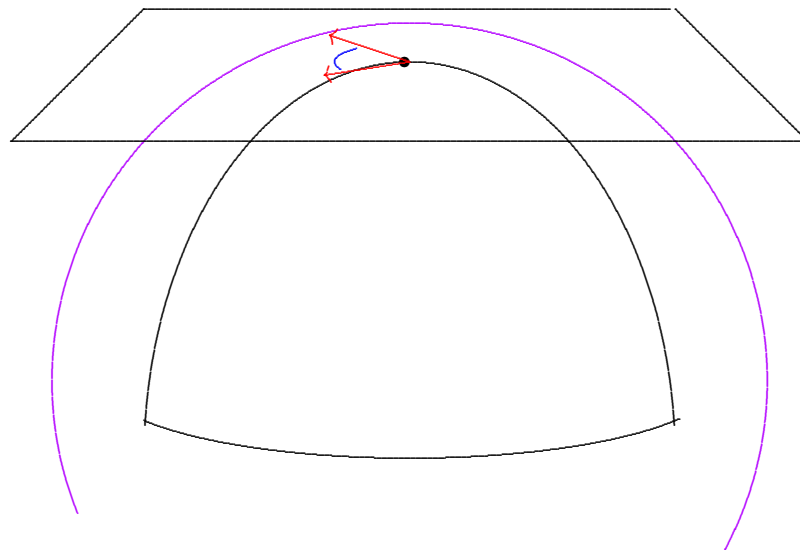
holonomy $\subset \mathbf{O}(n)$



Kähler metrics:

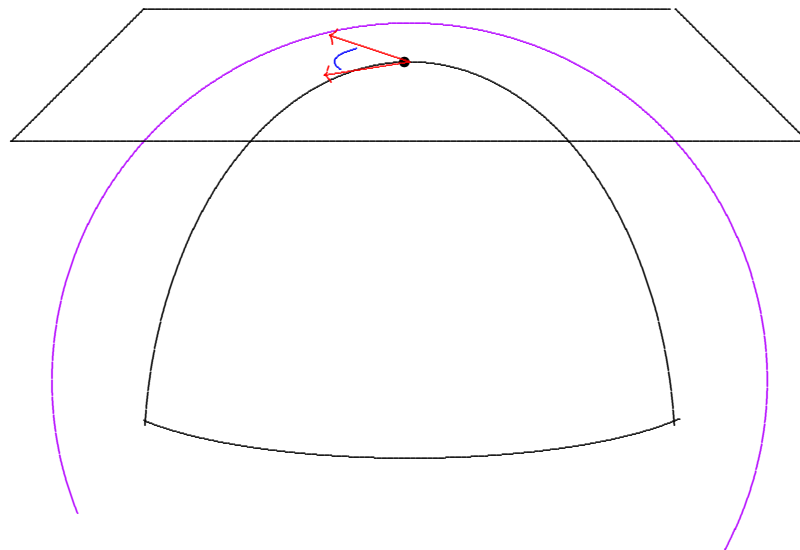
(M^{2m}, g) :

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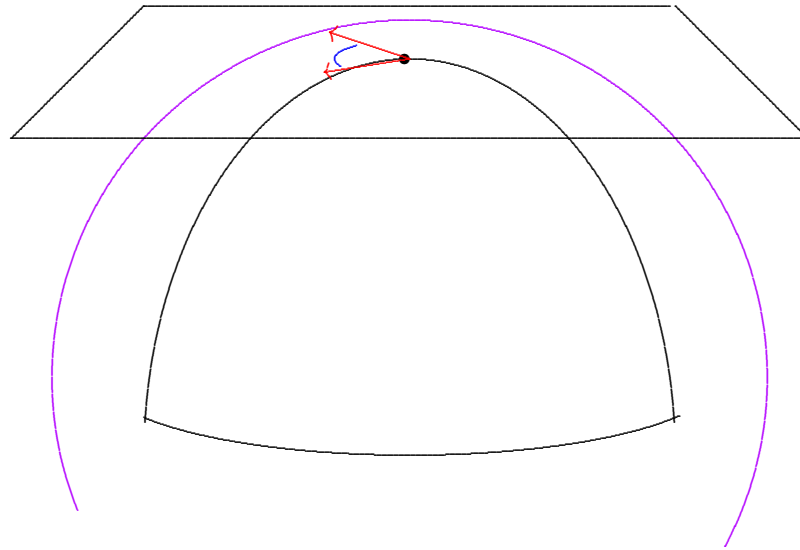
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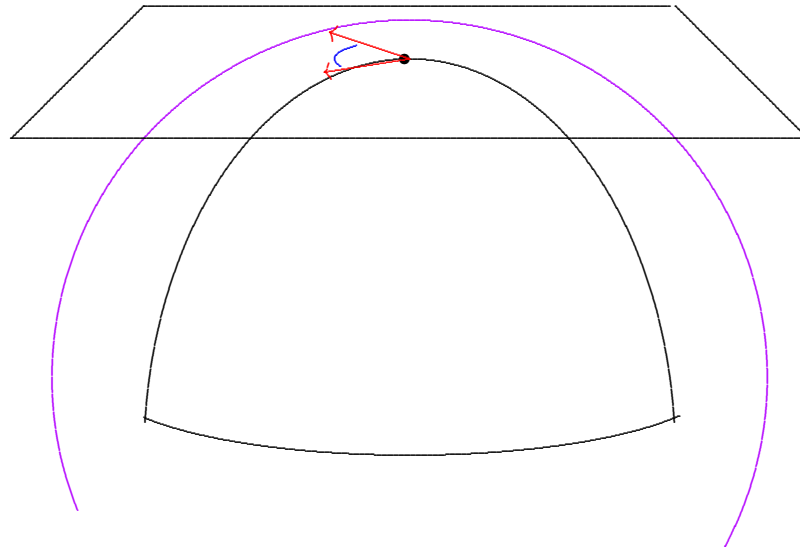
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$$\mathbf{U}(m) := \mathbf{O}(2m) \cap \mathbf{GL}(m, \mathbb{C})$$

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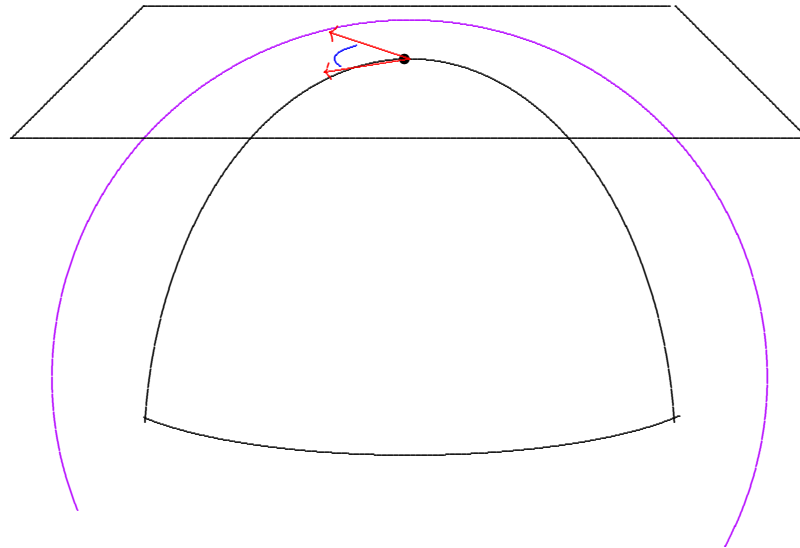
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Makes tangent space a complex vector space!

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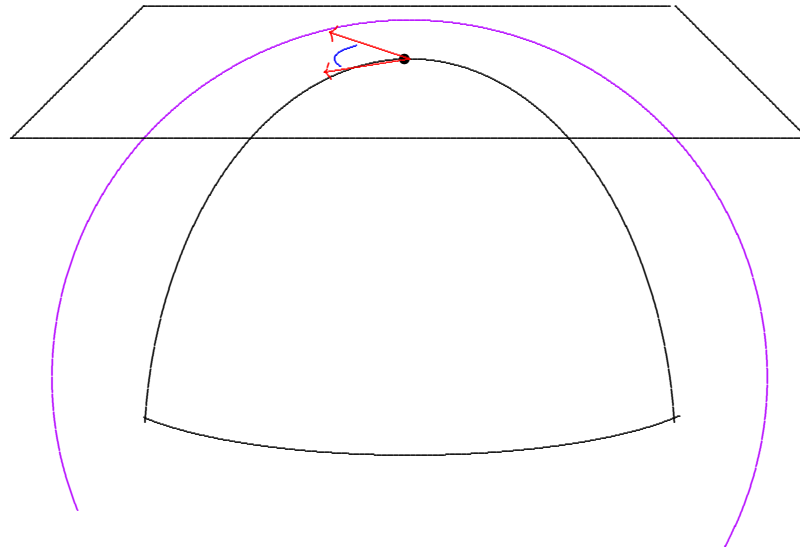
Makes tangent space a complex vector space!

$$J : TM \rightarrow TM, \quad J^2 = -\text{identity}$$

“almost-complex structure”

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Makes tangent space a complex vector space!

Invariant under parallel transport!

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ω called “Kähler form.”

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$$\omega = i \sum_{j,k=1}^m \frac{\partial^2 f}{\partial z^j \partial \bar{z}^k} dz^j \wedge d\bar{z}^k$$

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Kähler magic:

$$r = - \sum_{j,k=1}^m \frac{\partial^2}{\partial z^j \partial \bar{z}^k} \log \det[g_{p\bar{q}}] \left[dz^j \otimes d\bar{z}^k + d\bar{z}^k \otimes dz^j \right]$$

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If we define the Ricci form by

$$\rho = r(J\cdot, \cdot)$$

then $i\rho$ is curvature of canonical line bundle $\Lambda^{m,0}$.

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Kähler condition simplifies the Einstein condition!

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Hitchin: Every Einstein g on $K3$ is Calabi-Yau.

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By contrast, high-dimensional Einstein metrics too common; have little to do with geometrization.

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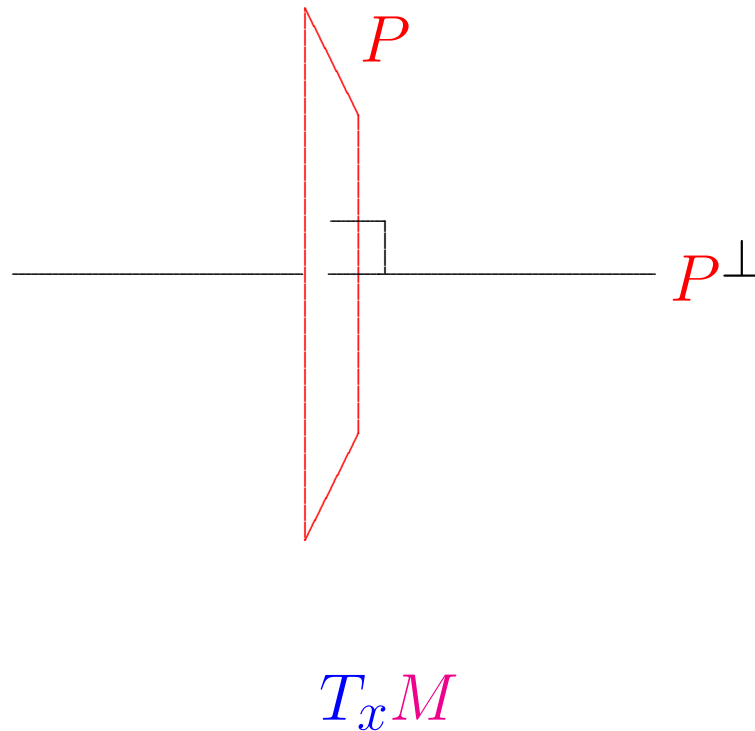
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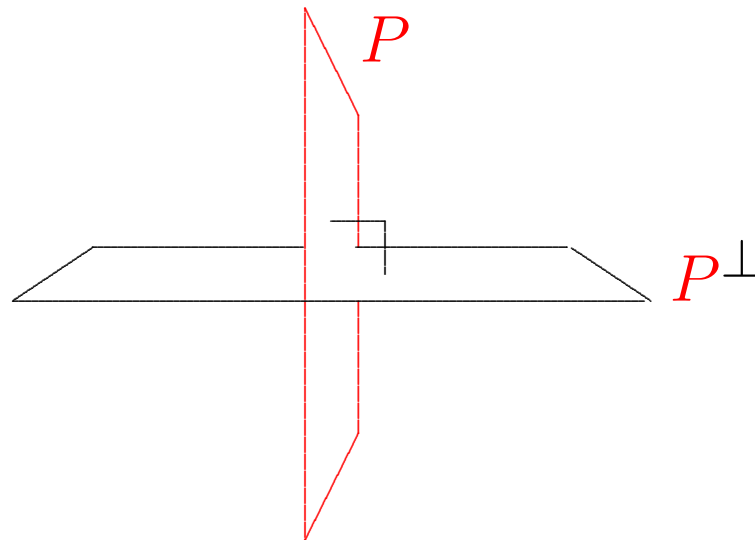
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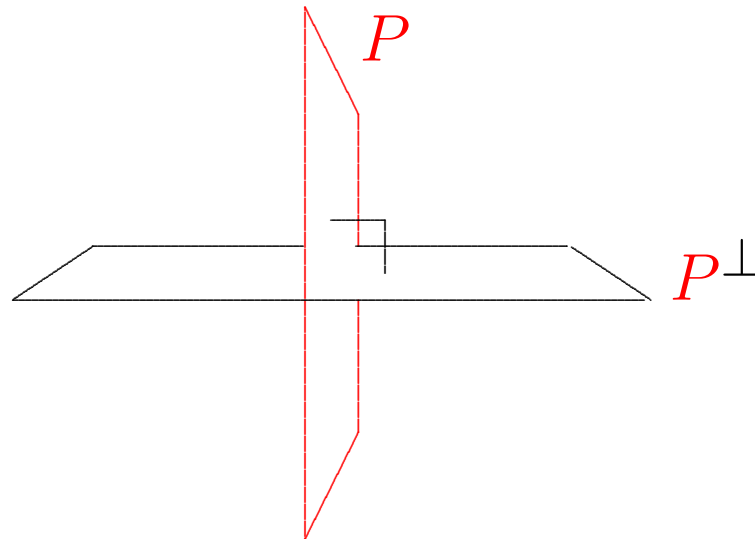


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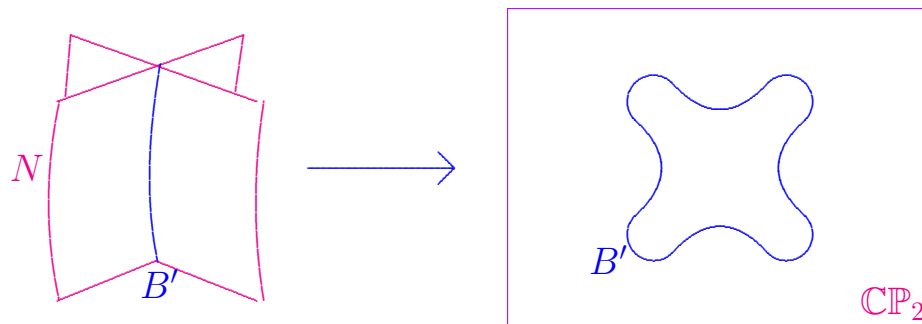
$$j : M \hookrightarrow \mathbb{C}\mathbb{P}_k$$

*such that $c_1(M)$ is negative multiple of $j^*c_1(\mathbb{C}\mathbb{P}_k)$.*

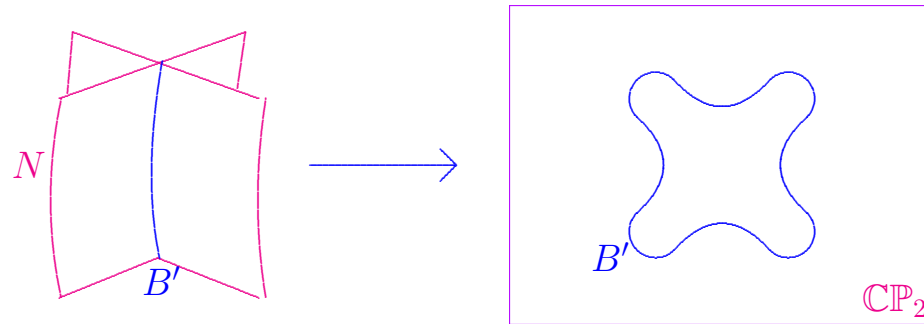
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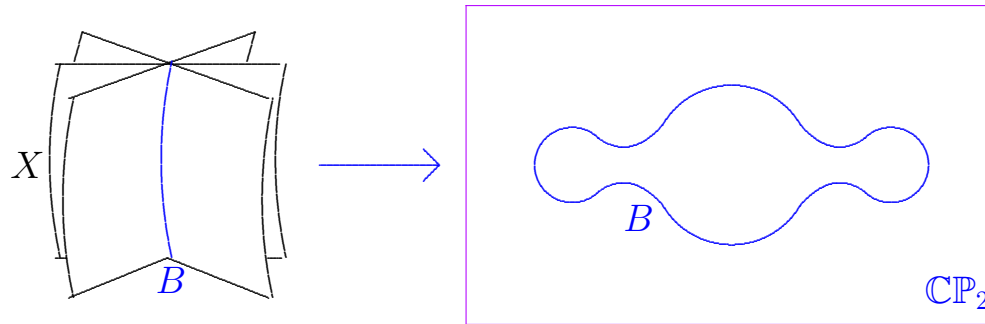


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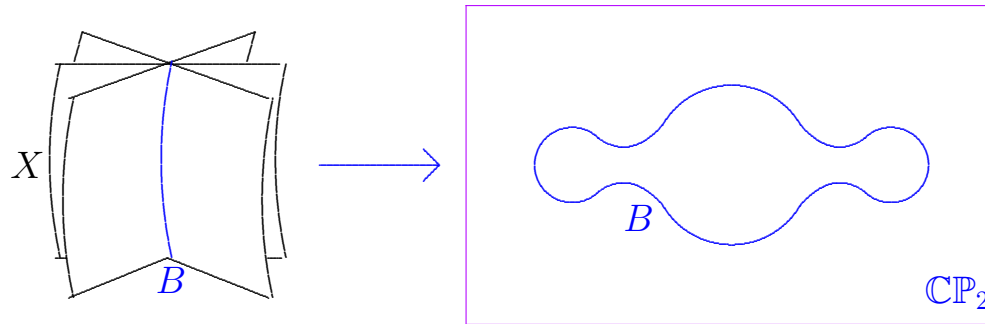


Aubin/Yau $\implies N$ carries Einstein metric.

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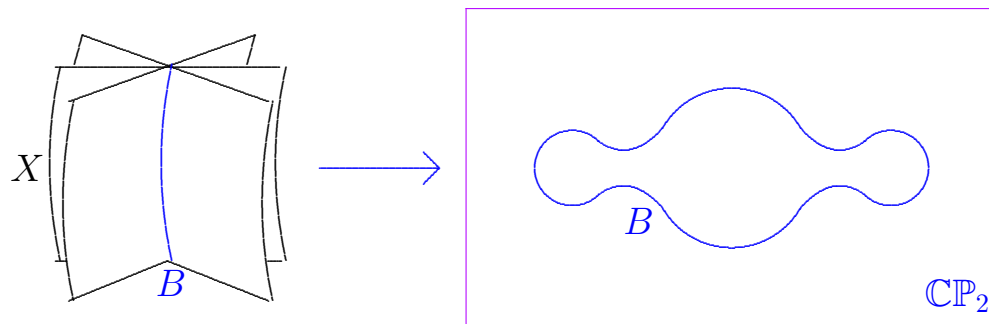
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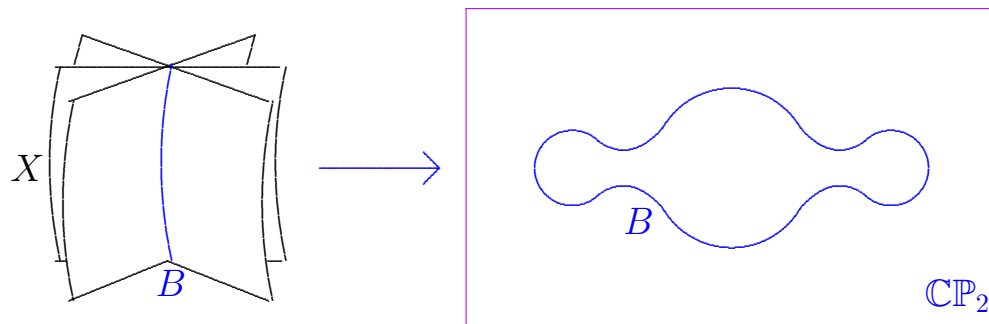


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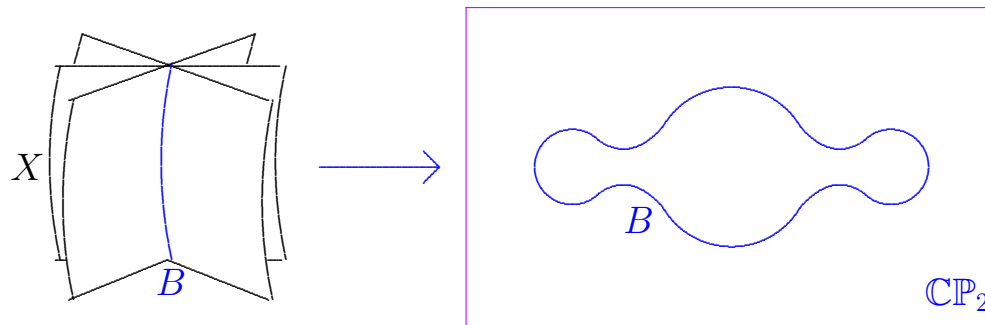
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Theorem \implies *no* Einstein metric on M .

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Moral: Existence depends on diffeotype!

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